

# Super-Resolution for Bright XMM-Newton Sources

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# What is Super-Resolution (SR)?



Super-resolution example from Google Brain Team's SR3  
(<https://arxiv.org/pdf/2104.07636.pdf>)

- Digital image processing technique to enhance the resolution of images beyond their original quality
- Approaches:
  - Classical techniques
  - Deep-learning based techniques
- Applications:
  - Medical imagery
  - Surveillance
  - Restoring pictures
  - Satellite images

# How is Super-Resolution Useful for Us?

- Super-resolution **cannot**
  - provide unambiguous output images
  - provide outputs that can be used for analyses such as spectra
- However, it **can**
  - learn common noise and distortion patterns
  - provide a perceptual improvement in image clarity
  - increase the confidence about the presence of certain features (e.g. for deblending)

Super-Resolution can serve as a visual explorative tool and provide inspiration and more confidence for new proposals!

## Deep Learning-Based Super-Resolution and De-Noising for XMM-Newton Images

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### ABSTRACT

The field of artificial intelligence based image enhancement has been rapidly evolving over the last few years and is able to produce impressive results on non-astronomical images. In this work we present the first application of Machine Learning based super-resolution (SR) and de-noising (DN) to enhance X-ray images from the European Space Agency's *XMM-Newton* telescope. Using *XMM-Newton* images in band [0.5,2] keV from the European Photon Imaging Camera pn detector (EPIC-pn), we develop *XMM-SuperRes* and *XMM-DeNoise* — deep learning-based models that can generate enhanced SR and DN images from real observations. The models are trained on realistic *XMM-Newton* simulations such that *XMM-SuperRes* will output images with two times smaller point-spread function and with improved noise characteristics. The *XMM-DeNoise* model is trained to produce images with 2.5× the input exposure time from 20 to 50 ks. When tested on real images, DN improves the image quality by 8.2%, as quantified by the global peak-signal-to-noise ratio. These enhanced images allow identification of features that are otherwise hard or impossible to perceive in the original or in filtered/smoothed images with traditional methods. We demonstrate the feasibility of using our deep learning models to enhance *XMM-Newton* X-ray images to increase their scientific value in a way that could benefit the legacy of the *XMM-Newton* archive.

- Super-Resolution Neural Network effectively increasing exposure time and resolution of XMM-Newton images
- Trained on real XMM-Newton images and 30855 simulated images generated with the SIXTE X-Ray simulation software
- Convolutional neural network with Residual-in-Residual Dense Block (RRDB)

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Metric	Input	Predicted ( <i>XMM-SuperRes</i> )
L1	0.01096	0.006508
PSNR	33.525	38.034
Poisson	0.08285	0.04997
SSIM	0.8248	0.907
MS_SSIM	0.9499	0.9846
FSIM	0.8657	0.8688
HaarPsi	0.5312	0.697

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# XMM-SuperRes: Network Architecture

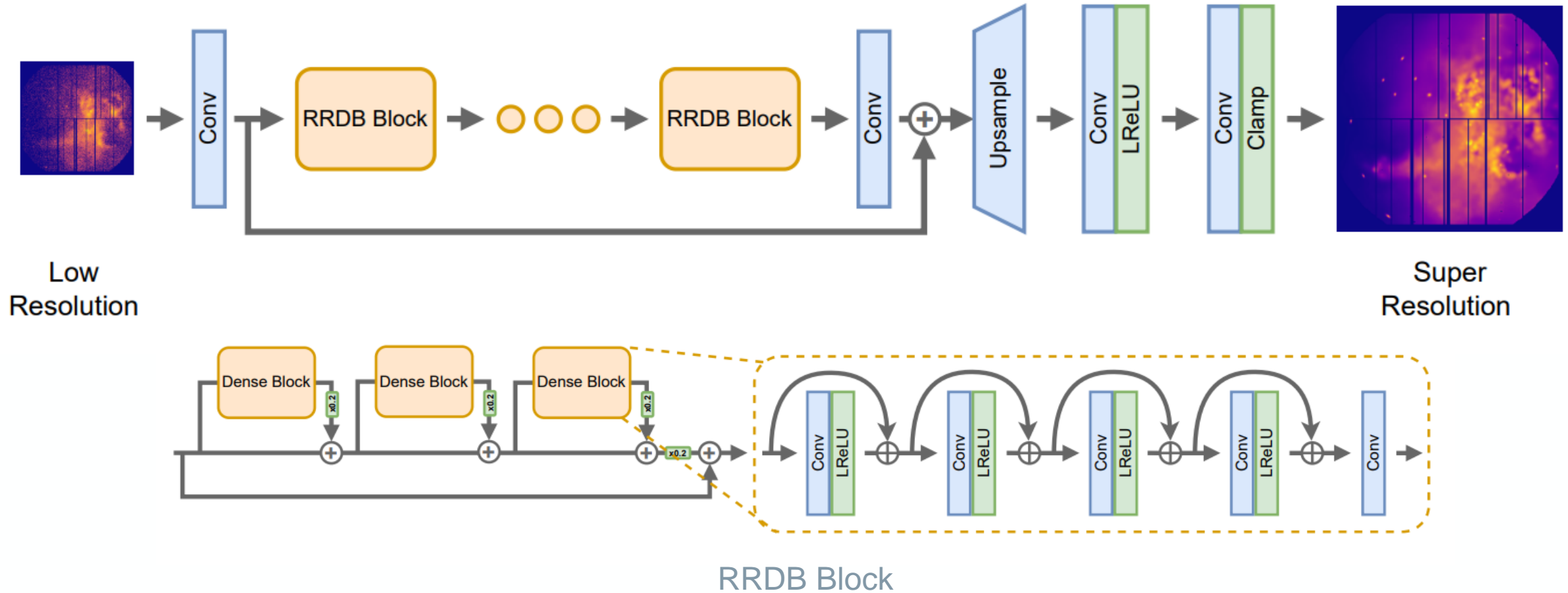
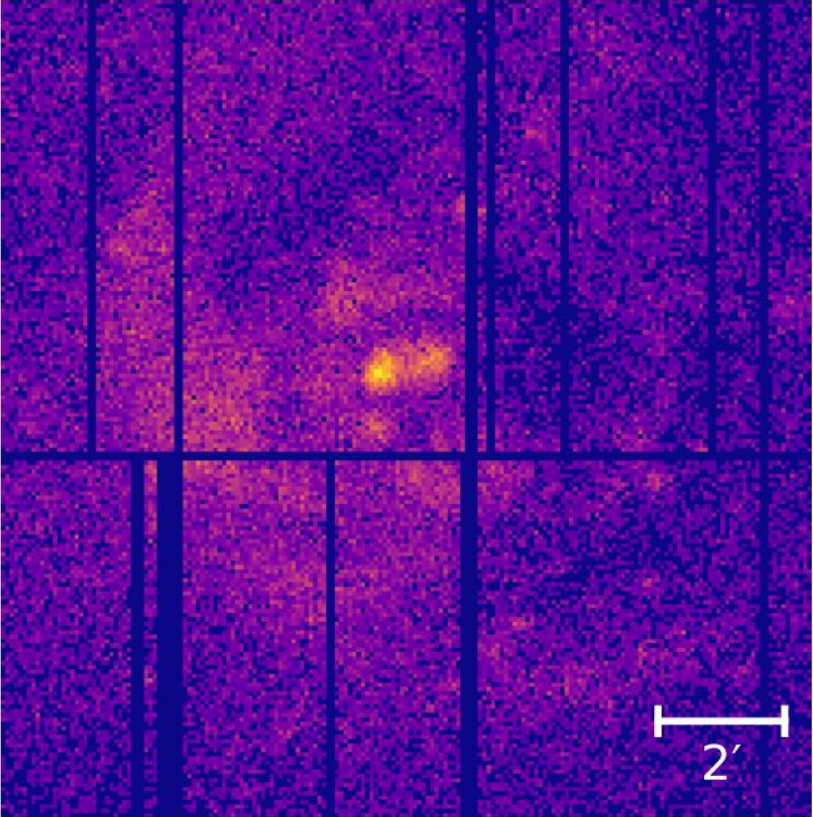


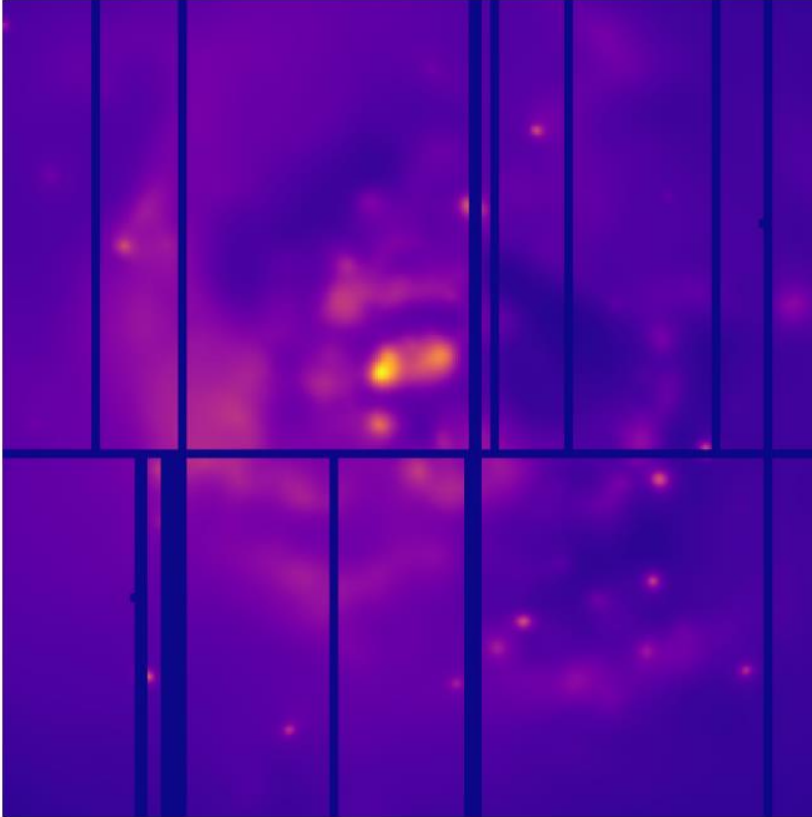
Image from <https://arxiv.org/pdf/2205.01152.pdf>

# XMM-SuperRes Example

Input



Prediction



Label

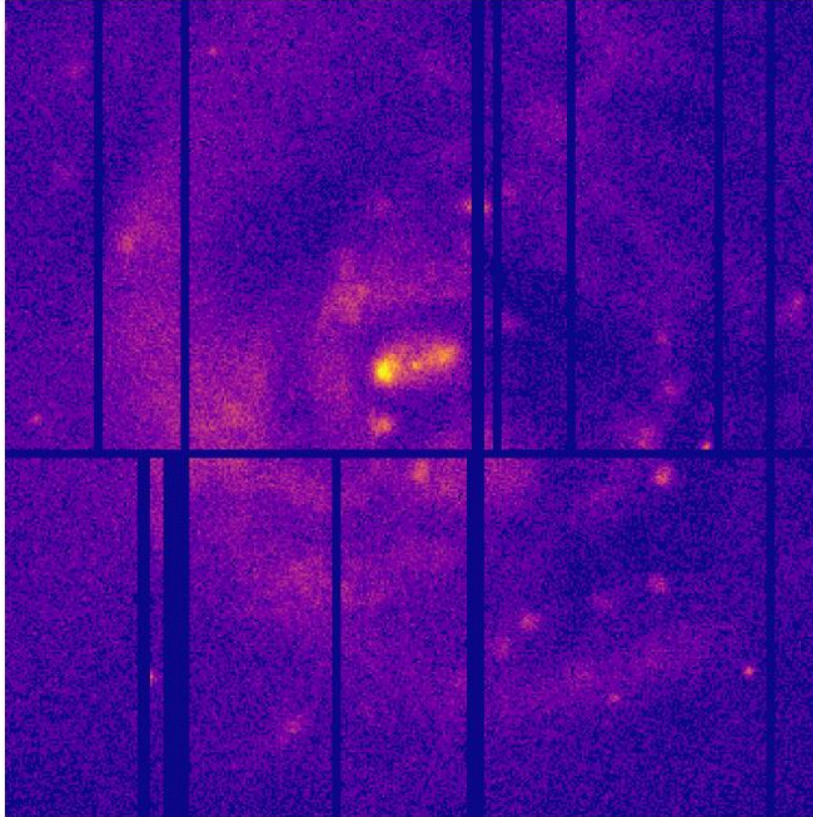


Image from <https://arxiv.org/pdf/2205.01152.pdf>

# XMM-SuperRes Example

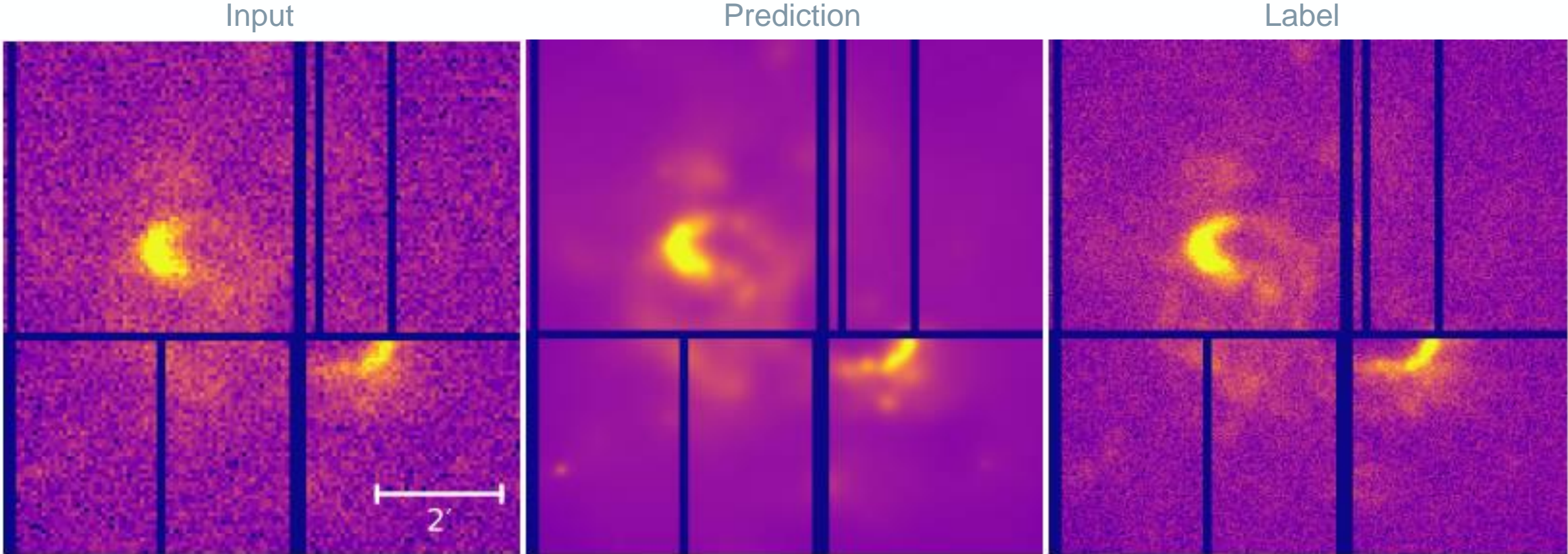


Image from <https://arxiv.org/pdf/2205.01152.pdf>

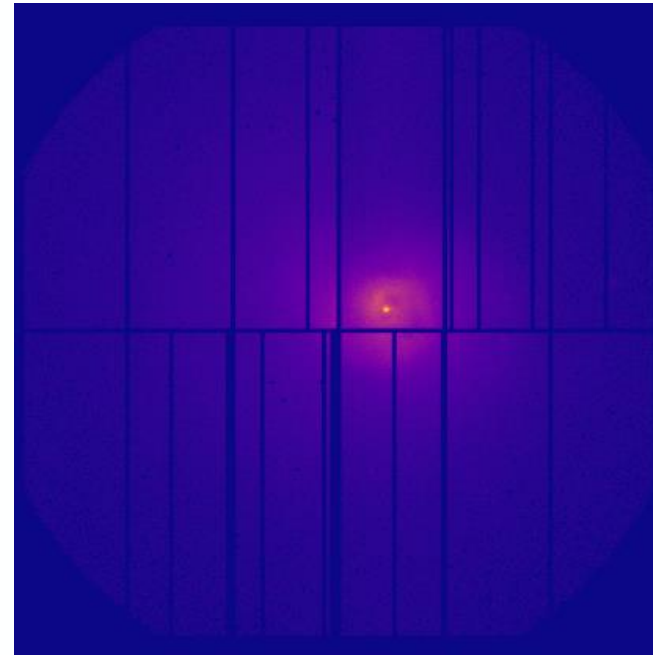


The current implementation of XMM-SuperRes shows two major limitations:

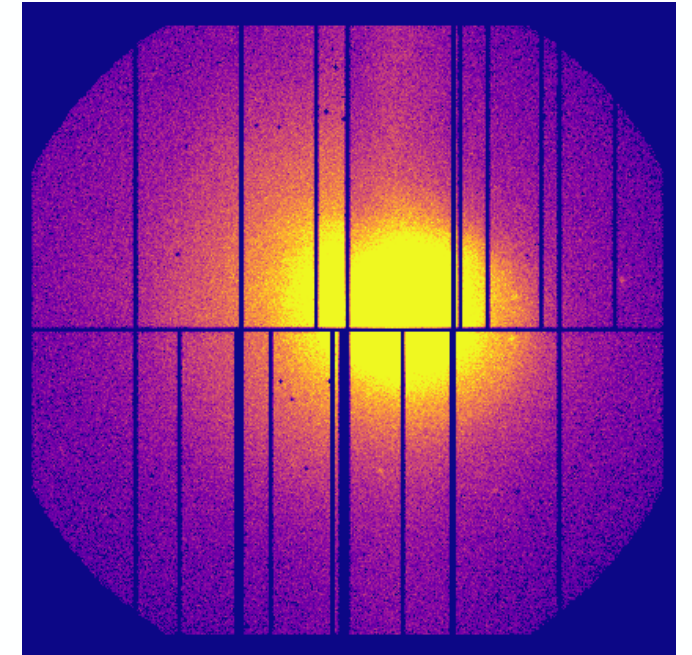
1. The currently employed clamping leads to a loss of information in very bright regions

2. While a reduction in error has been shown, it is quite small and the ability to super-resolve has not been explicitly tested

- XMM-SuperRes employs clamping to avoid instability during training
- A constant value is used, defined based on the background noise
- **Due to the doubled resolution, the target clamping value is defined as a quarter of the input clamping value**
- However, using this constant value can lead to the loss of valuable information for bright sources

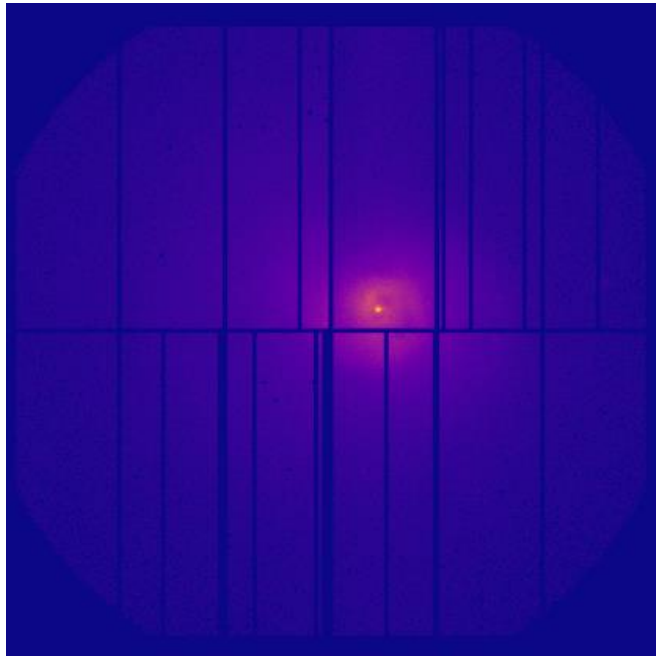


Unclamped input

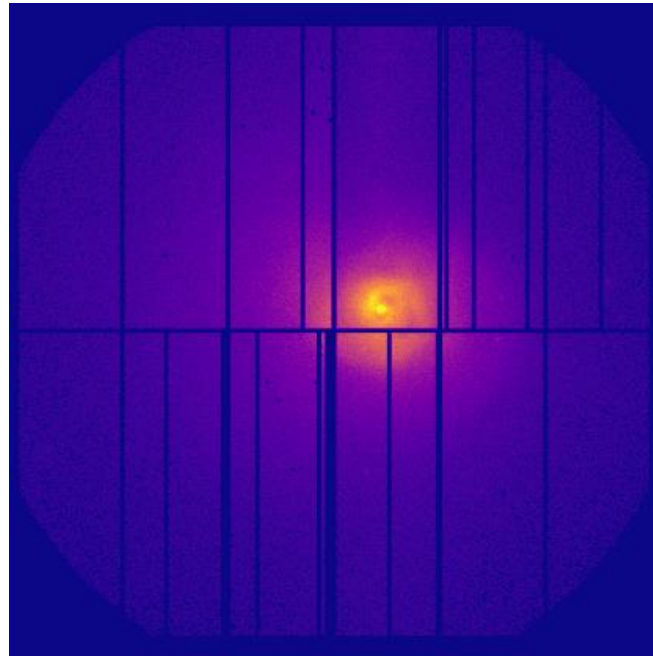


Clamped input

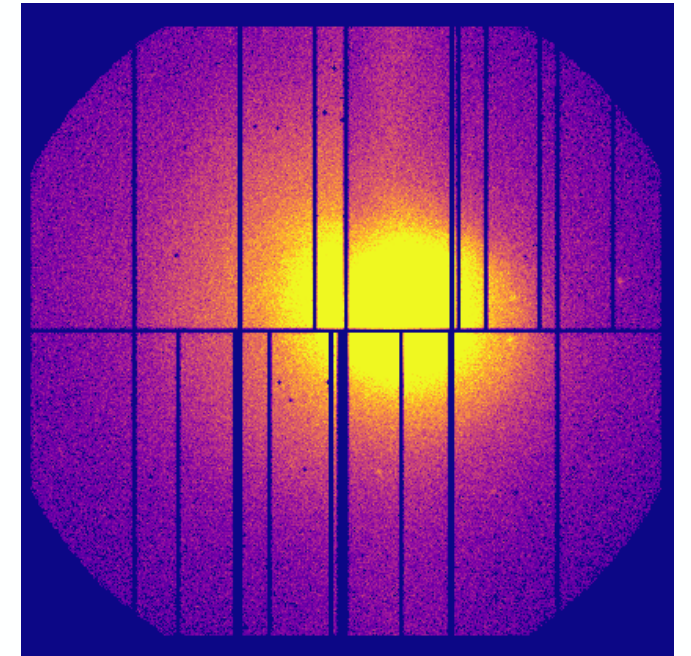
- Replaced constant clamping by adaptive quantile clamping
- Even when choosing a quantile of 0.9999 the training was still stable
- However, the accuracy dropped to 0.8823 (MS-SSIM)/ 23.1563 (PSNR) ) (compared to 0.9846 (MS-SSIM)/ 38.034 (PSNR))



Unclamped



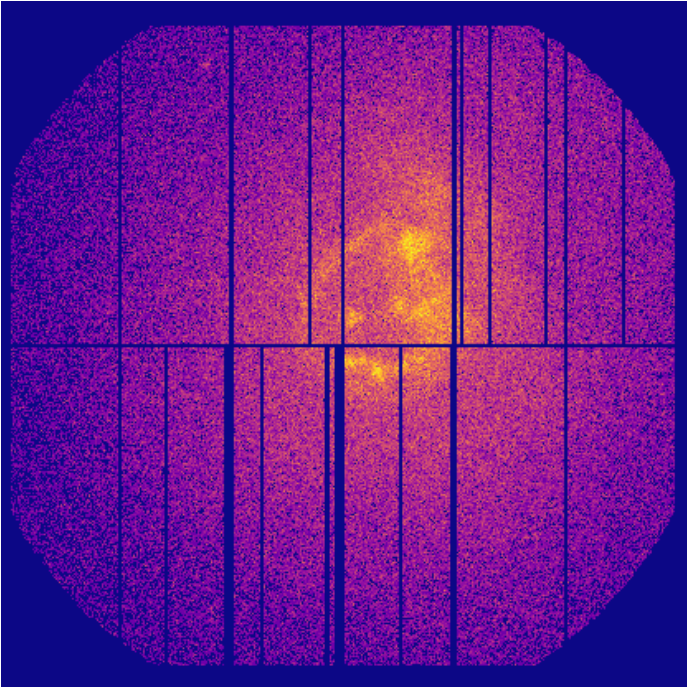
0.9999 quantile clamping



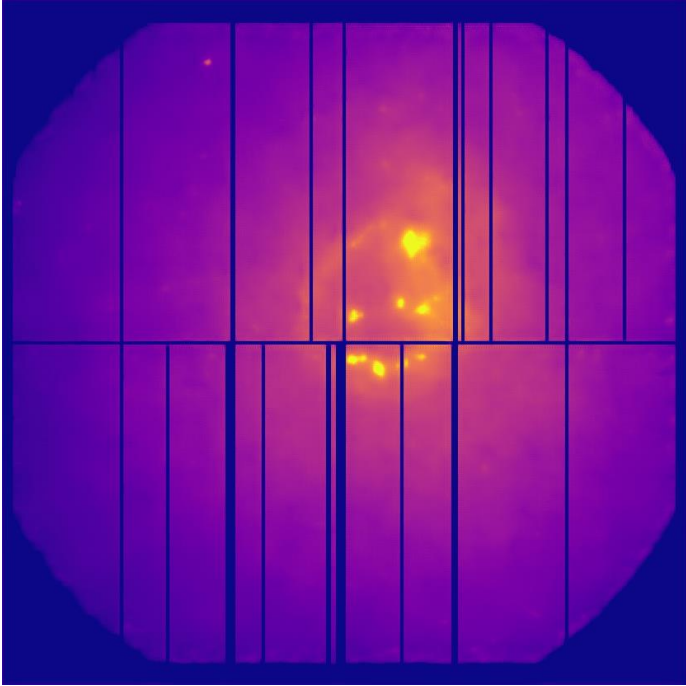
Constant clamping

# Quantile Clamping Results

Input



Prediction



Label

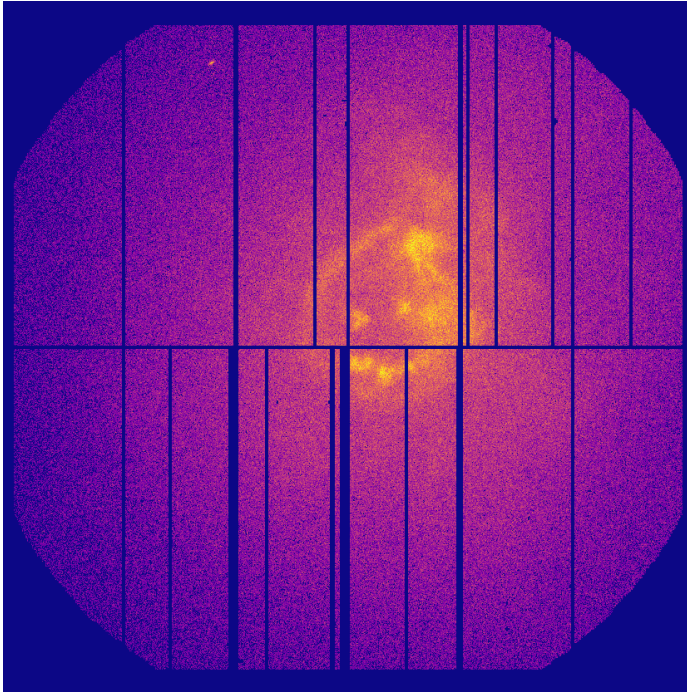
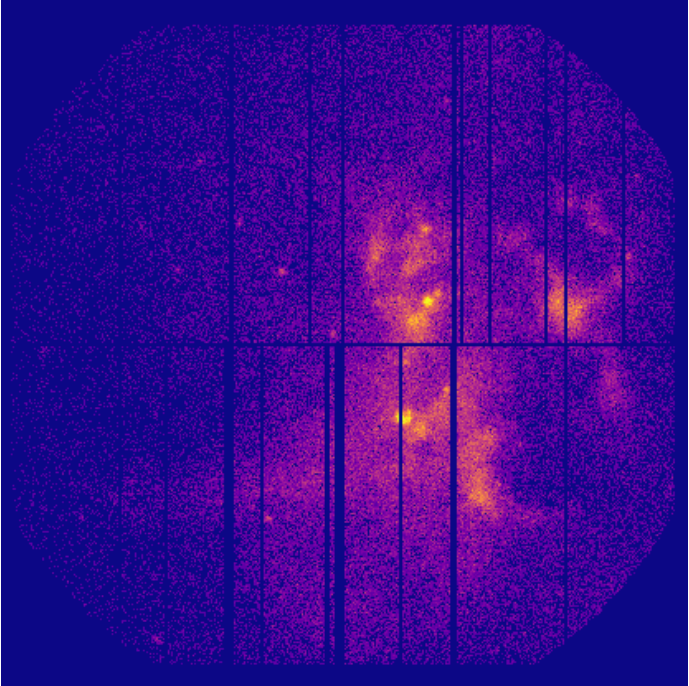


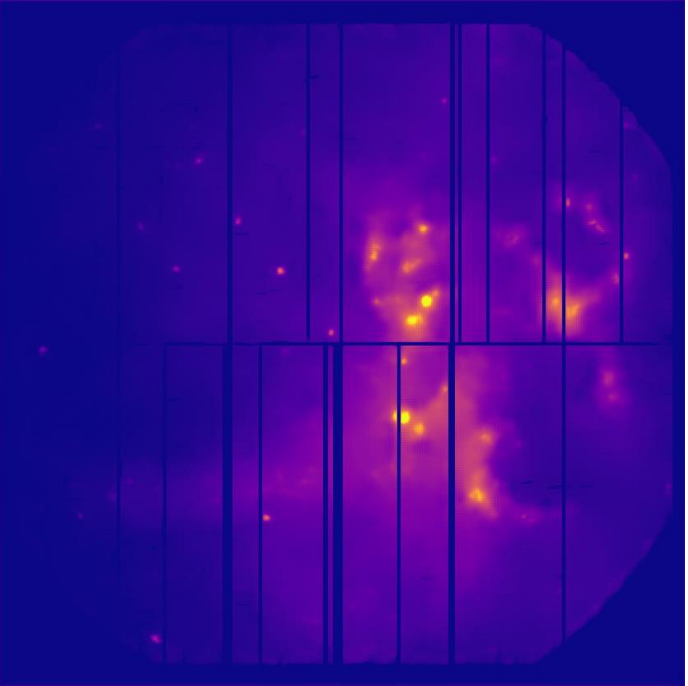
Image from <https://arxiv.org/pdf/2205.01152.pdf>

# Quantile Clamping Results

Input



Prediction



Label

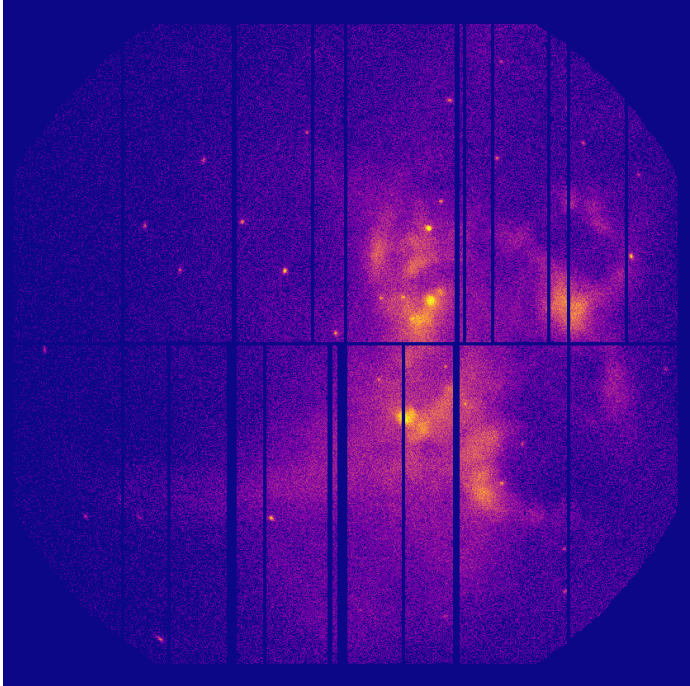
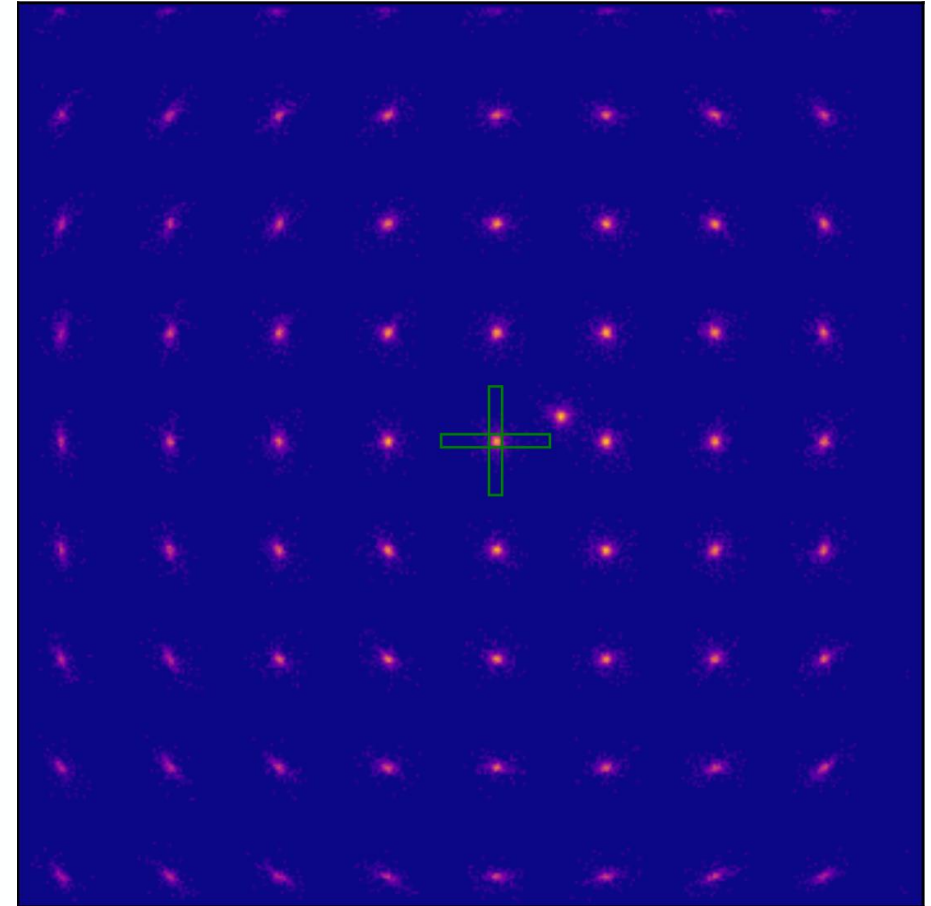


Image from <https://arxiv.org/pdf/2205.01152.pdf>

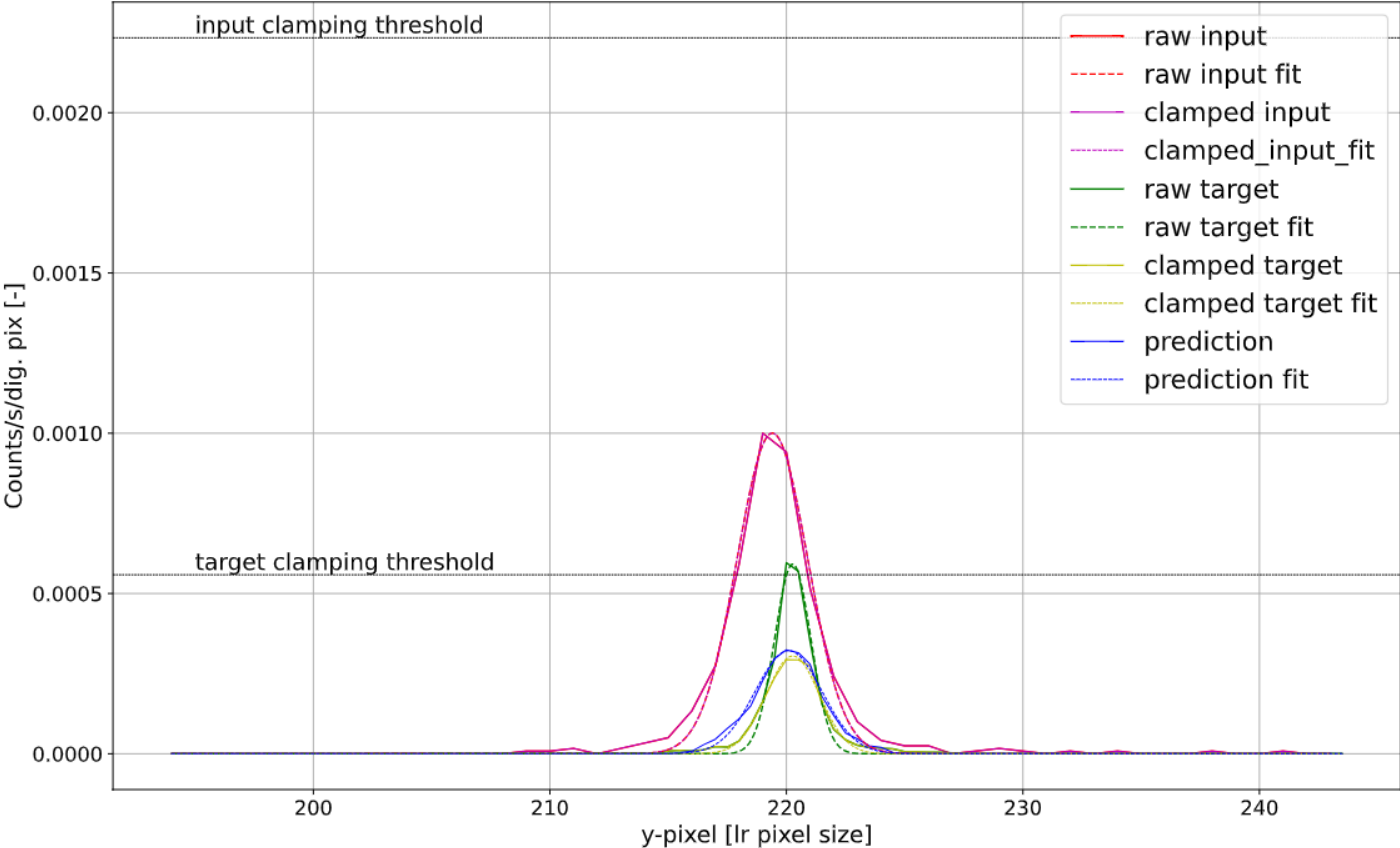
- The width of the point spread function is proportional to the resolution of the image
- Used simulated grid of point sources as input to XMM-SuperRes
- Summed up the counts within a small window of both the input and output
- Fit Gaussians to the results to see if the width decreases



# Testing the Point Spread Function

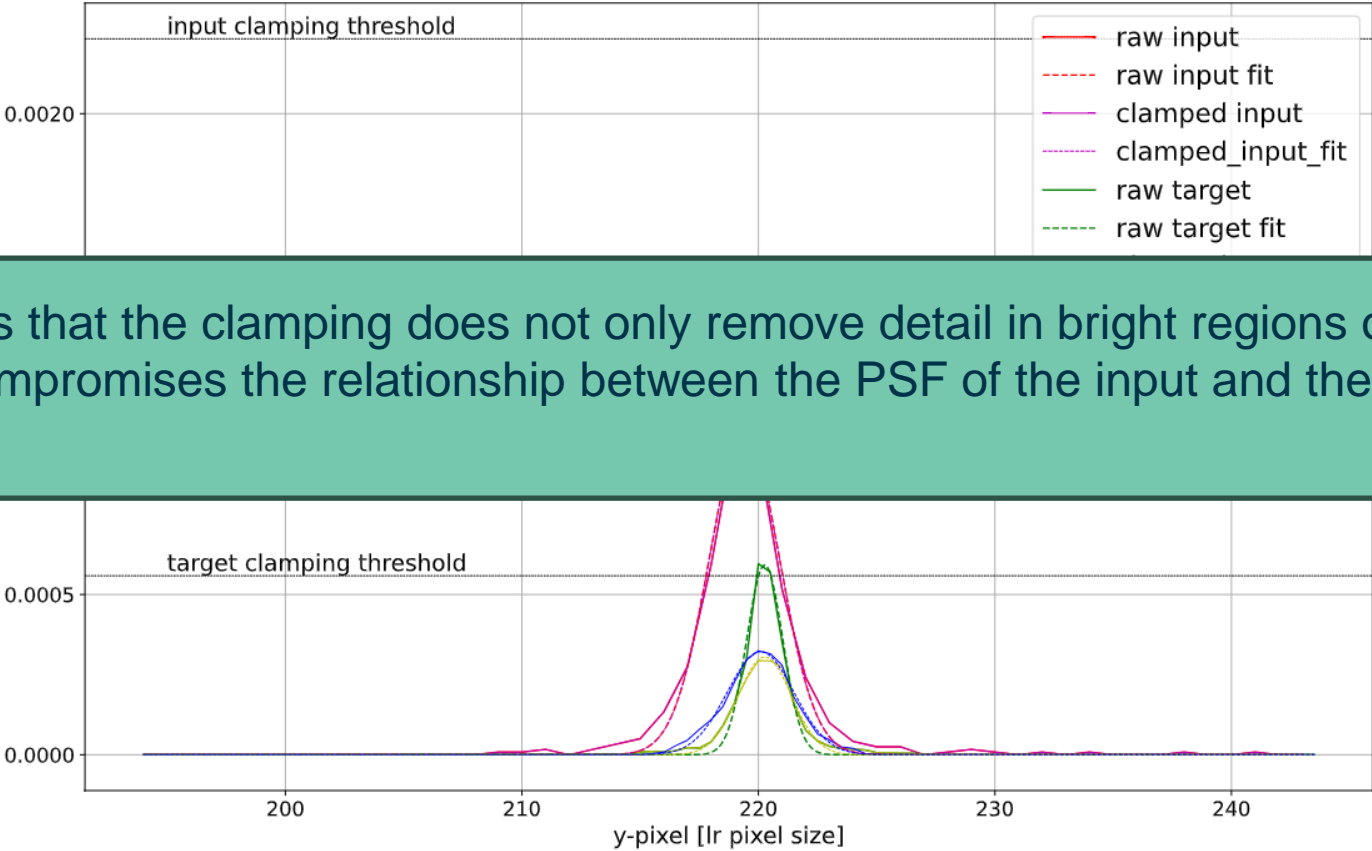


Input Fit Std = 1.50  
Clamped Input Fit Std = 1.50  
Target Fit Std = 0.77  
Clamped Target Fit Std = 1.13  
Prediction Fit Std = 1.41



# Testing the Point Spread Function

Input Fit Std = 1.50  
Clamped Input Fit Std = 1.50  
Target Fit Std = 0.77  
Clamped Target Fit Std = 1.13  
Prediction Fit Std = 1.41



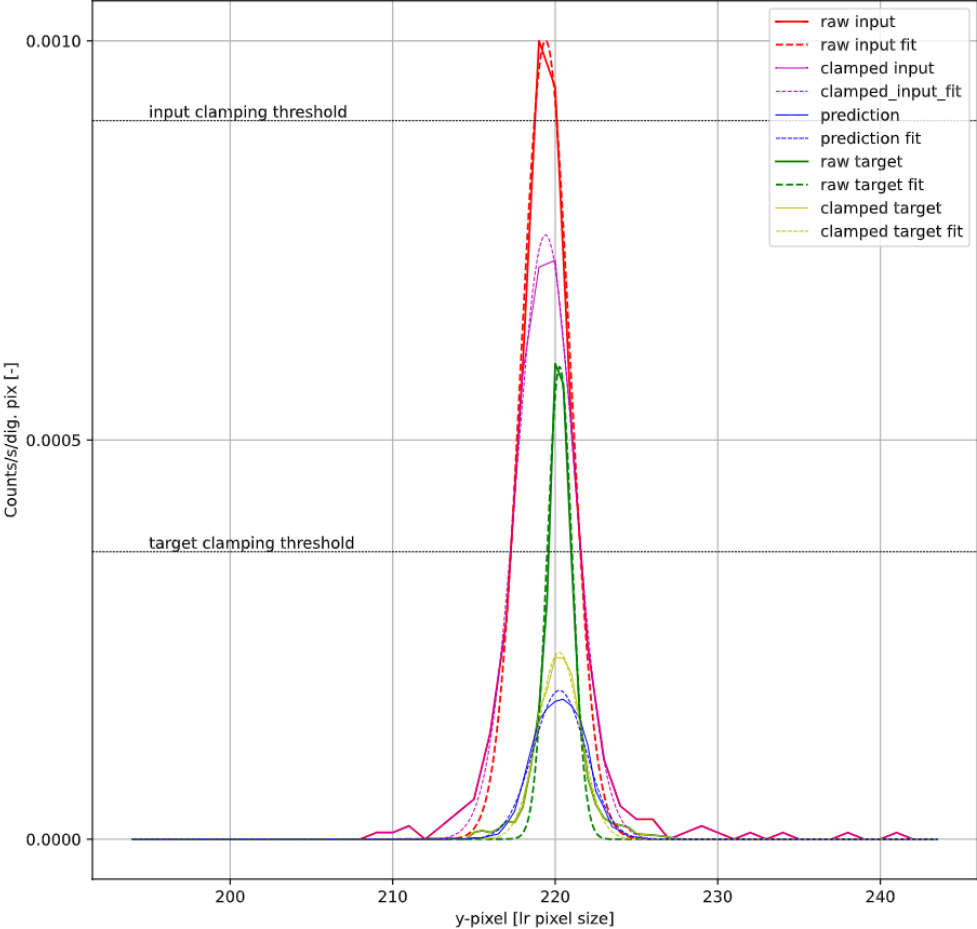
This implies that the clamping does not only remove detail in bright regions of the image, but also compromises the relationship between the PSF of the input and the target!



# Testing the PSF with Quantile Clamping

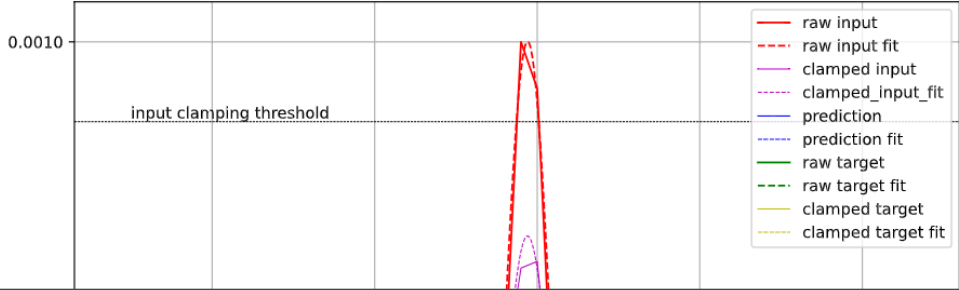


Input Fit: Amplitude = 0.0060, Mean = 219.41, Std = 1.50  
Clamped Input Fit: Amplitude = 0.0045, Mean = 219.40, Std = 1.78  
Target Fit: Amplitude = 0.01, Mean = 220.27, Std = 0.77  
Clamped Target Fit: Amplitude = 0.0028, Mean = 220.25, Std = 1.28  
Prediction Fit: Amplitude = 0.0022, Mean = 220.26, Std = 1.61

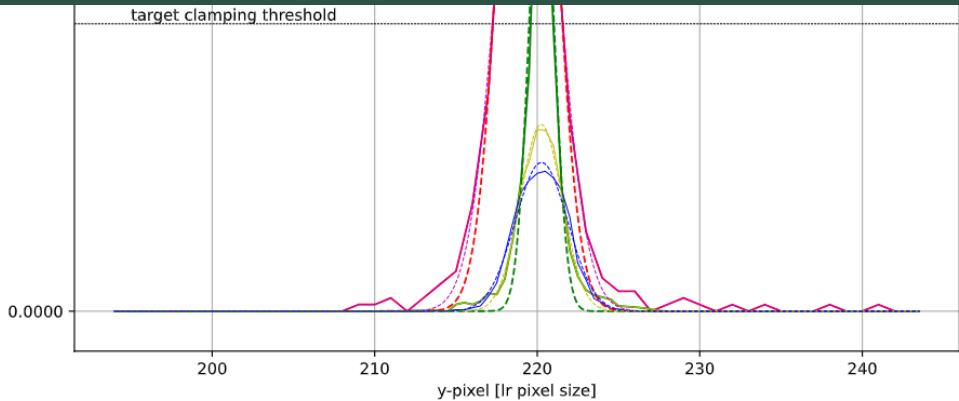


# Testing the PSF with Quantile Clamping

Input Fit: Amplitude = 0.0060, Mean = 219.41, Std = 1.50  
Clamped Input Fit: Amplitude = 0.0045, Mean = 219.40, Std = 1.78  
Target Fit: Amplitude = 0.01, Mean = 220.27, Std = 0.77  
Clamped Target Fit: Amplitude = 0.0028, Mean = 220.25, Std = 1.28  
Prediction Fit: Amplitude = 0.0022, Mean = 220.26, Std = 1.61



Even with quantile clamping, the relationship between the PSF of the input and target is compromised!



- Super-Resolution can serve as a visual explorative tool and provide inspiration and more confidence for new proposals
- It already provides output images with perceived increase in resolution and decrease in noise
- Using quantile clamping overcomes the issue of lost detail in bright regions
- However, both constant and quantile clamping compromise the relationship between the input and target PSF
- The high dynamic range of X-ray images makes the application of super-resolution tricky, and the clamping needs to be chosen with care!
  
- Outlook:
  - Look into ways to adjust the network such that no clamping is required
  - Adjust occurrence of bright sources within the dataset
  - Possibly apply the approach to large parts of the catalogue