

A catalogue of photometric redshifts for 4XMM

Angel Ruiz (IAASARS/NOA)

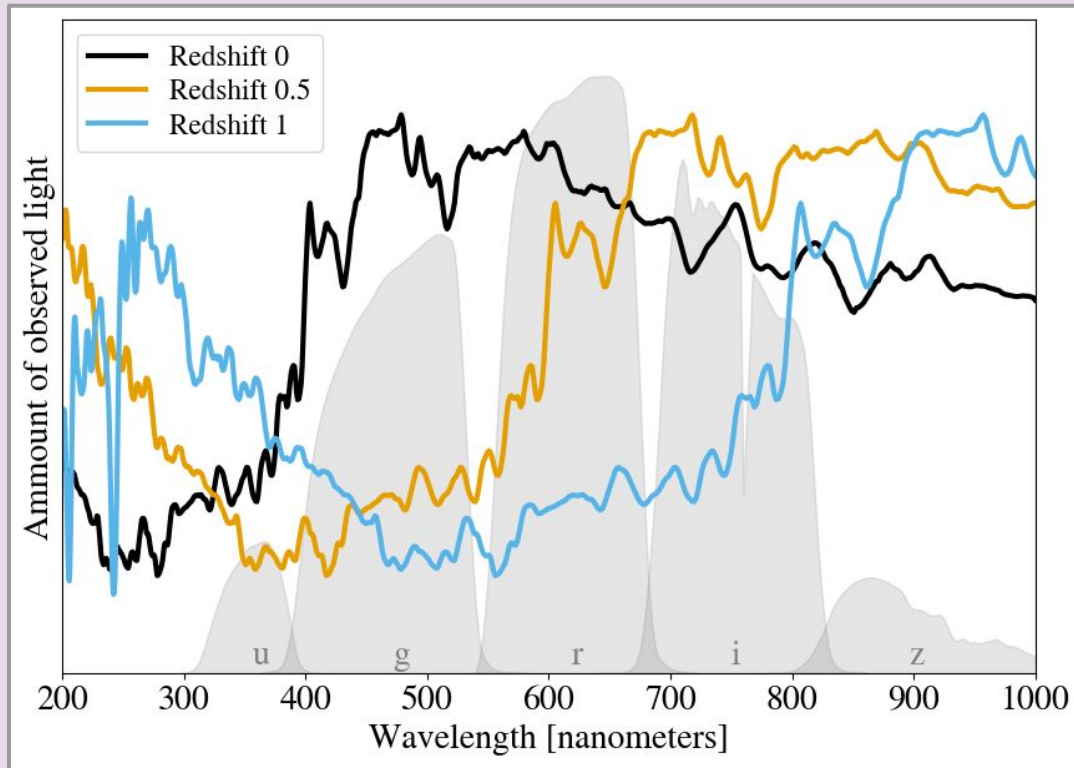
**E. Pouliaxis, I. Georgantopoulos,
A. Akylas, A. Georgakakis
& XMM2ATHENA team.**

INTRODUCTION

- The 4XMM serendipitous catalogue (**Webb+2020**) is one of the largest X-ray catalogues currently available, with ~630,000 unique sources in data release 12.
- In order to fully untap its scientific potential, distance estimations (i.e. redshifts) are needed.
- Spectroscopic redshifts are available for only a small fraction, and they are observationally very expensive.
- We need to rely on **photometric redshifts**.

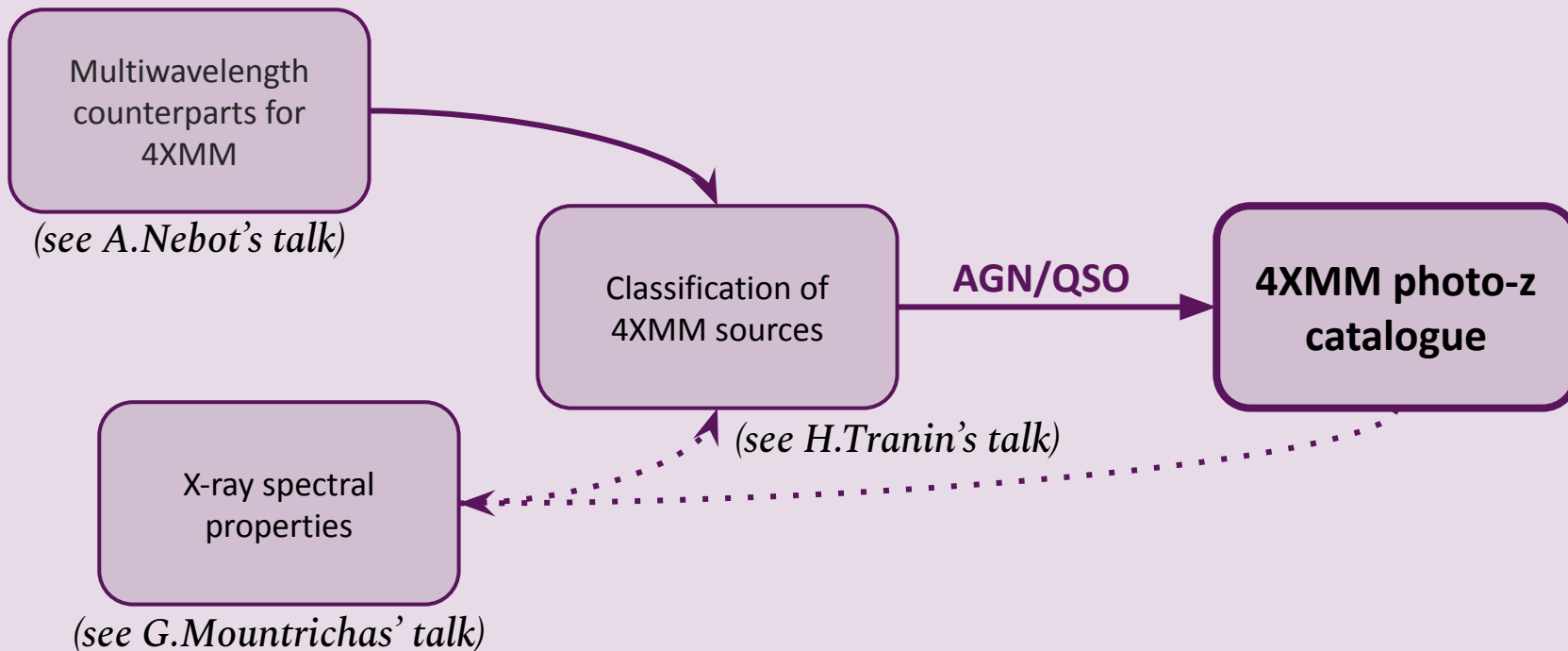


PHOTOMETRIC REDSHIFTS



- Observed colors for a given galaxy depend on the redshift (and galaxy type)
- Two major techniques for photo-z:
 - Template fitting
 - Supervised machine learning

XMM2ATHENA: Photo-z for 4XMM

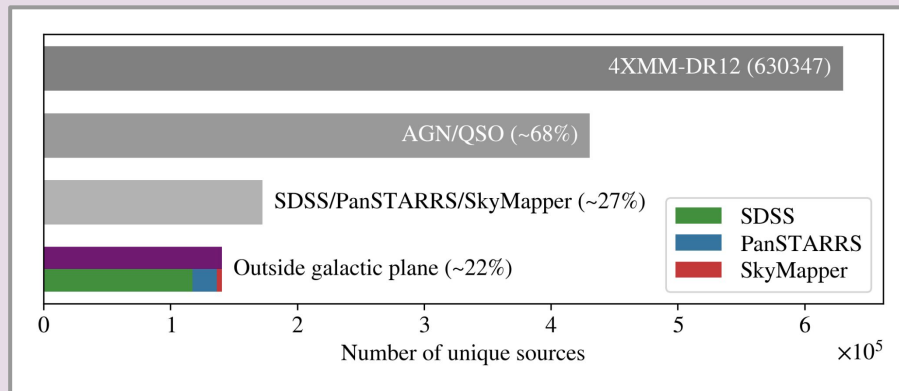


Development of techniques and tools for 5XMM and future Athena serendipitous catalogues.

■ 4XMM-DR12

(full sky, $\sim 1400 \text{ deg}^2$)

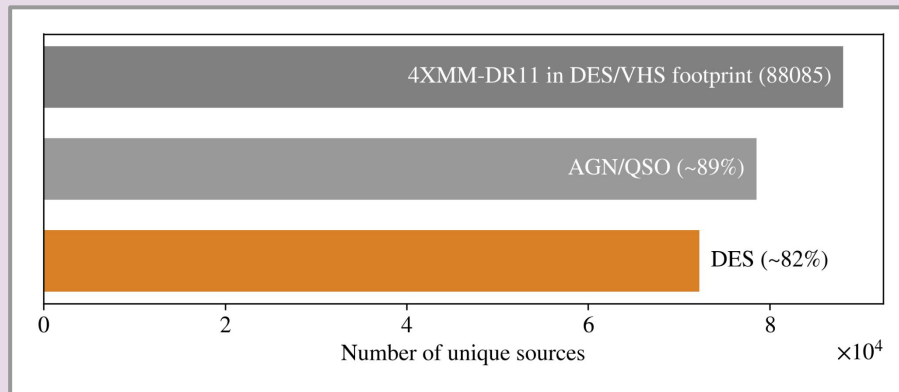
- Classified as AGNs
- Counterparts in SDSS/PanSTARRS/SkyMapper
- Outside the galactic plane



■ 4XMM-DR11

(inside DES/VHS footprint only, $\sim 200 \text{ deg}^2$)

- Classified as AGNs
- Counterparts in DES



We used two different photo-z techniques previously used for X-ray selected sources:

- **MLZ-TPZ** (Carrasco-Kind & Brunner 2013)
 - Supervised machine learning algorithm: **Random forest**
 - Used previously for **XXL-North** (Mountrichas+2017) and **3XMM** (XMMPZCAT, Ruiz+2018)
 - Training sample based on the MORX catalogue
 - Input attributes: **colors** and **optical extension**.

- **LePhare** (Arnouts+1997, Ilbert+2006)
 - SED template fitting code
 - Broadly used in X-ray surveys: **COSMOS** (Salvato+2009, 2011), **CDFS** (Hsu+2014), **Stripe82-X** (Ananna+2017), **eFEDS** (Salvato+2022), **XXL-North** (Pouliasis+2024)
 - Galaxy and AGN templates from eFEDS (Salvato+2022)

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Very good for the bulk of the population.
Needs rich photometry.

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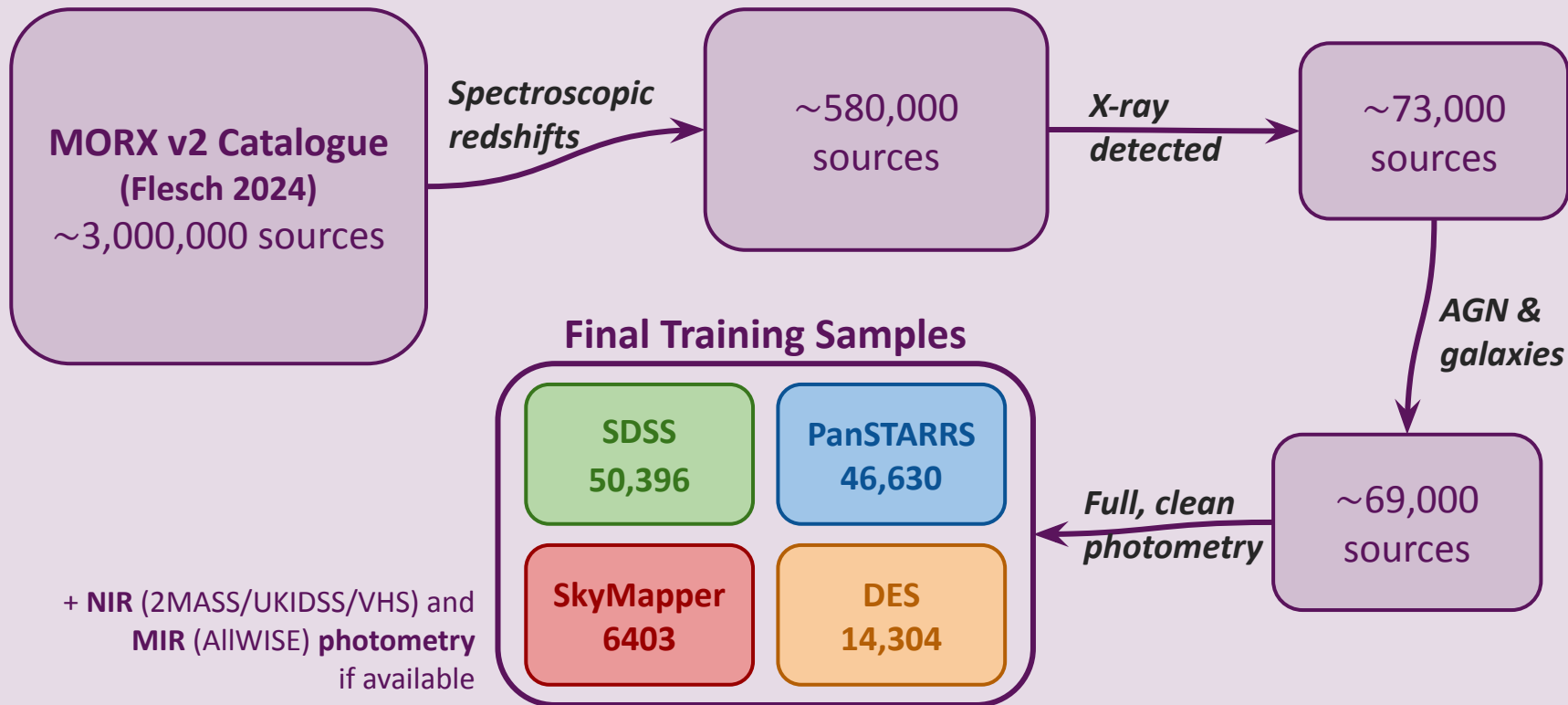
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(see E.Pouliasis' talk)

Incomplete photometry.
Rare sources
(e.g. high redshift).

XMM2ATHENA: Training/Validation Sample



XMM2ATHENA: Training/Validation Sample

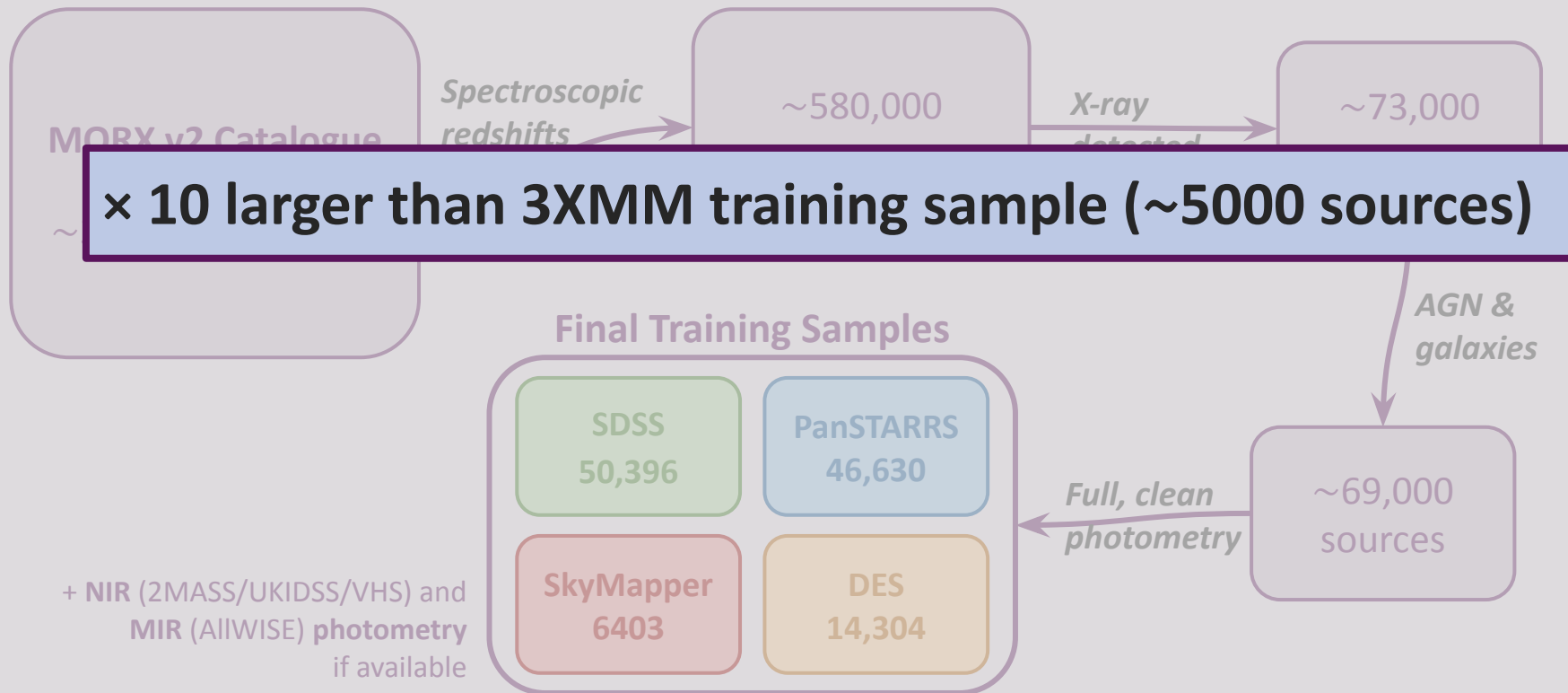
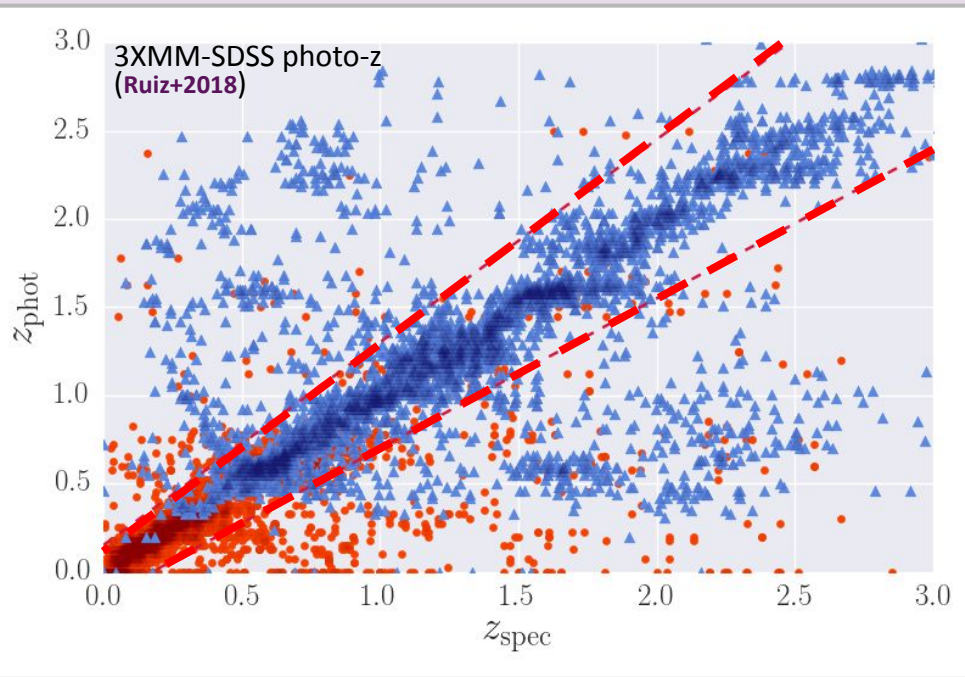


PHOTO-Z RELIABILITY AND ACCURACY



accuracy

$$\sigma_{\text{NMAD}} = 1.4826 \times \text{Median}(|\Delta z_{\text{norm}}|)$$

reliability

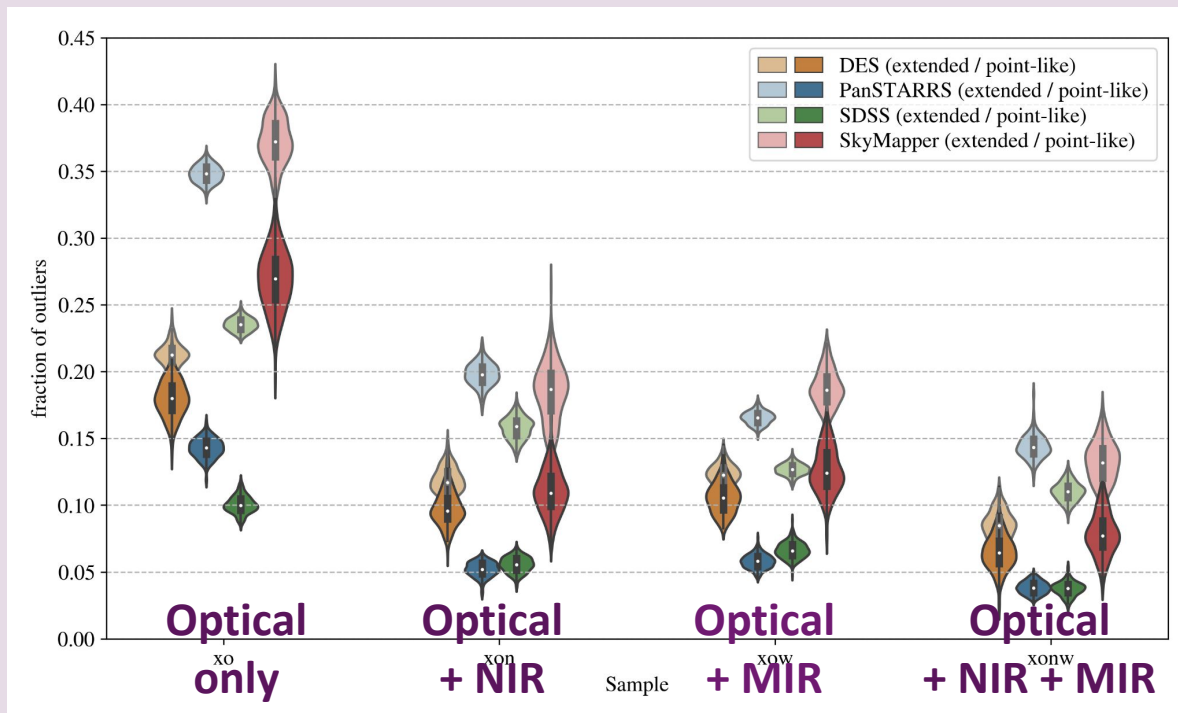
$$\eta = 100 \times N_{\text{outliers}} / N$$

$$\Delta z_{\text{norm}} = (z_{\text{spec}} - z_{\text{phot}}) / (1 + z_{\text{spec}})$$

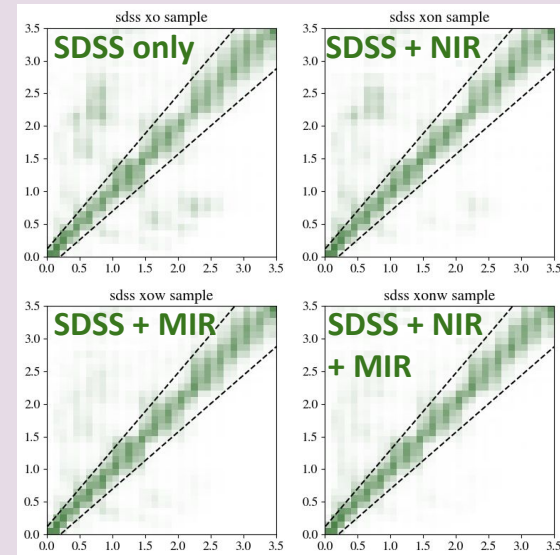
outlier if $|\Delta z_{\text{norm}}| > 0.15$

Results for η using TPZ

(6-fold cross-validation using the training samples)

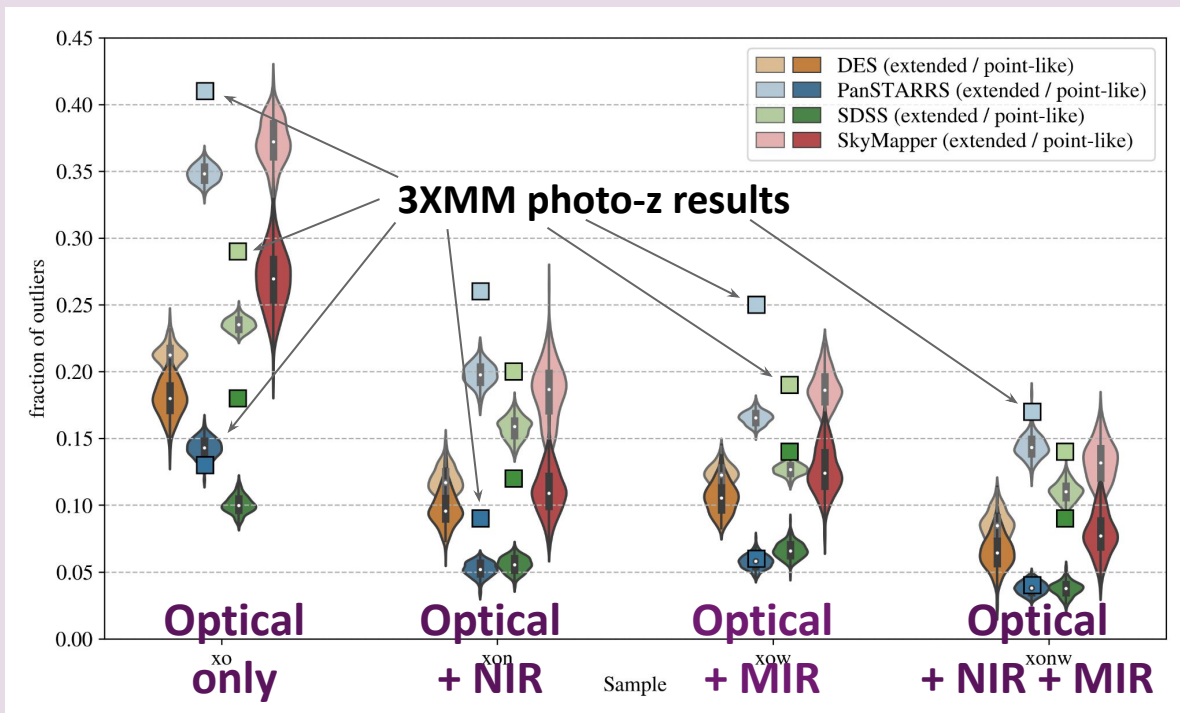


TPZ photo-z versus spec-z

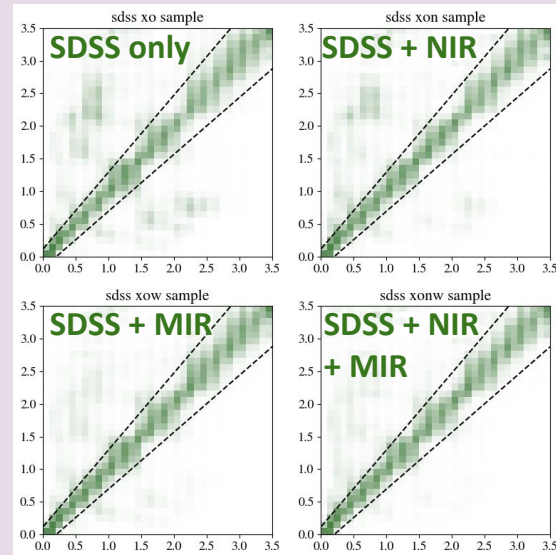


Results for η using TPZ

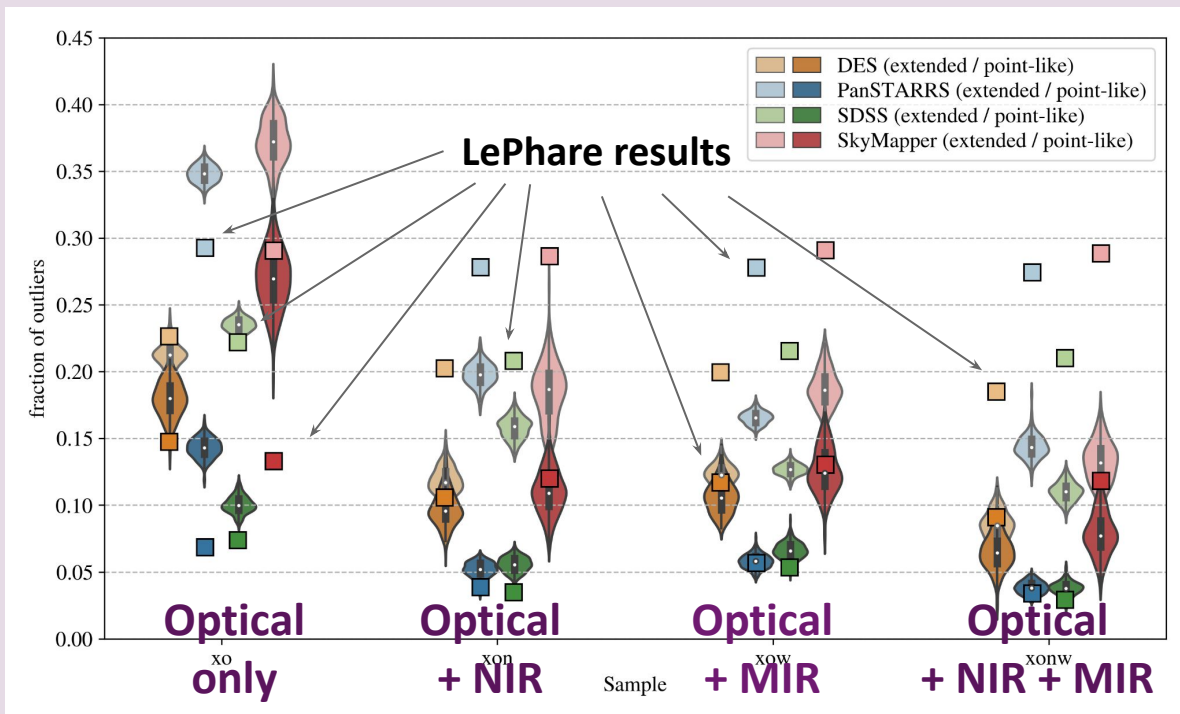
(6-fold cross-validation using the training samples)



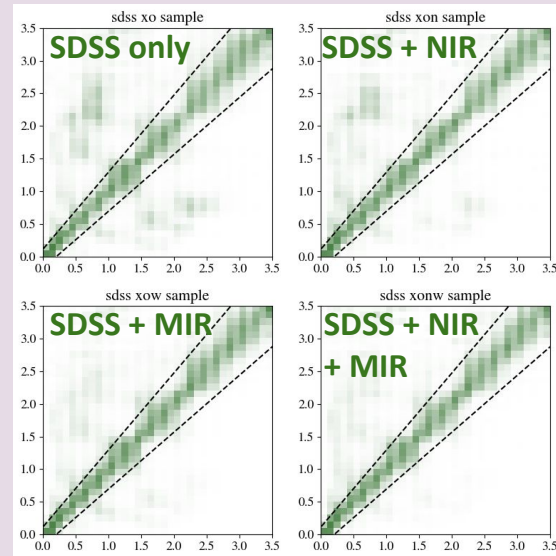
TPZ photo-z versus spec-z



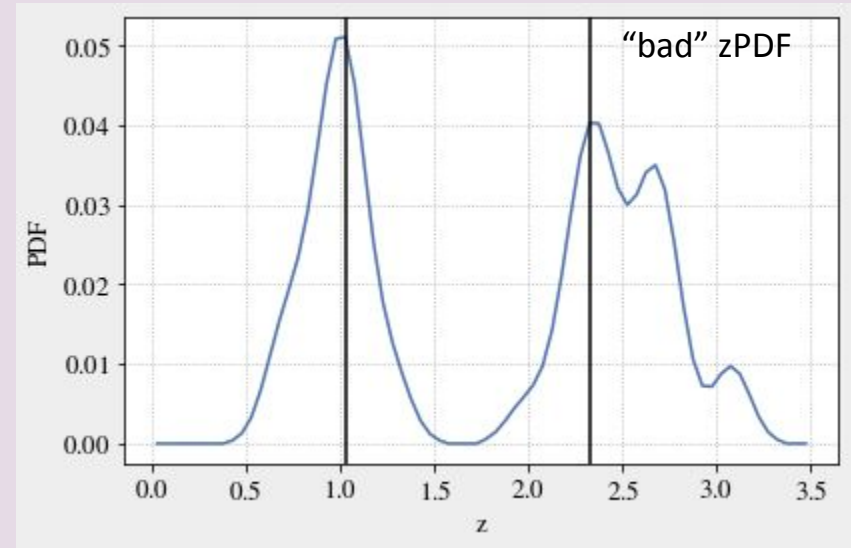
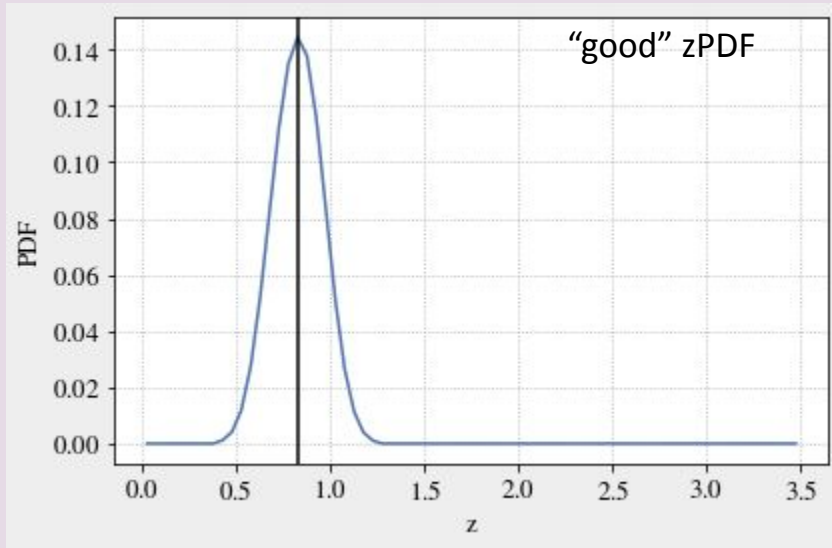
Results for η using LePhare (cross-validation using the training samples)



TPZ photo-z versus spec-z



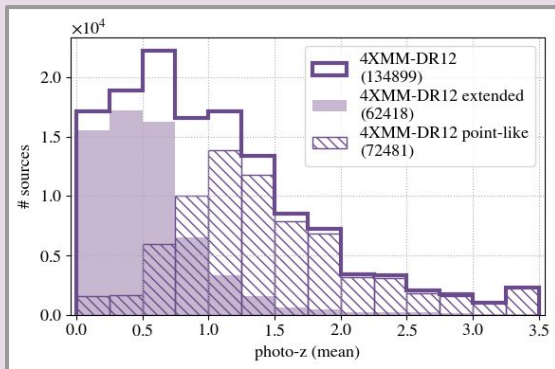
Reliability of TPZ and LePhare photo-z estimated through photometry coverage and the shape of the PDF: zConf, relative peak strength (PS) and number of peaks (NP) [as in Ruiz+2018].



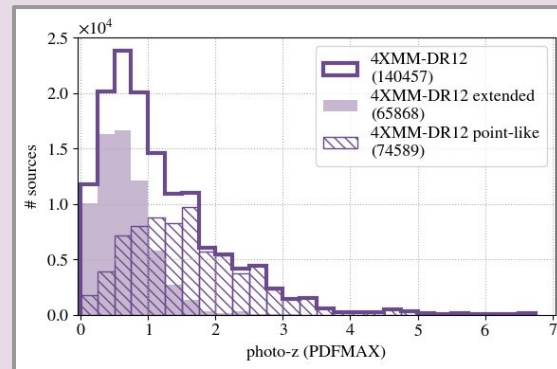
XMM2ATHENA: Photo-z results

4XMM-DR12
 ~1000 deg²

TPZ photo-z



LePhare photo-z

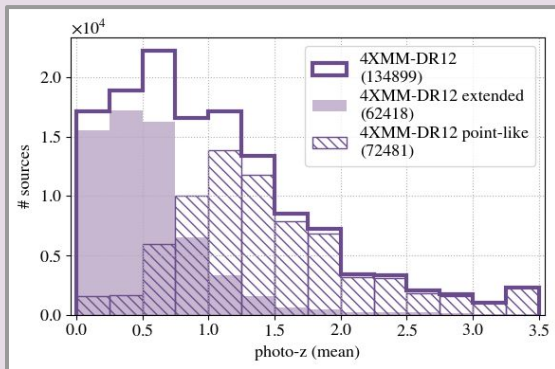


XMM2ATHENA: Photo-z results

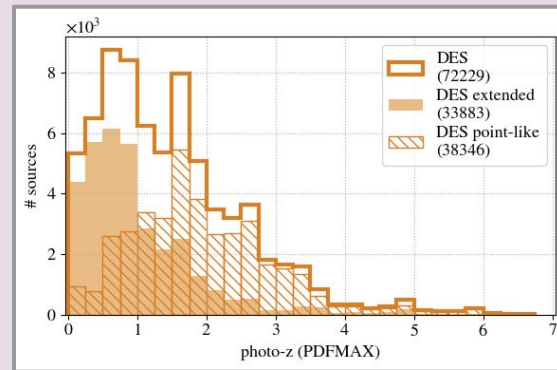
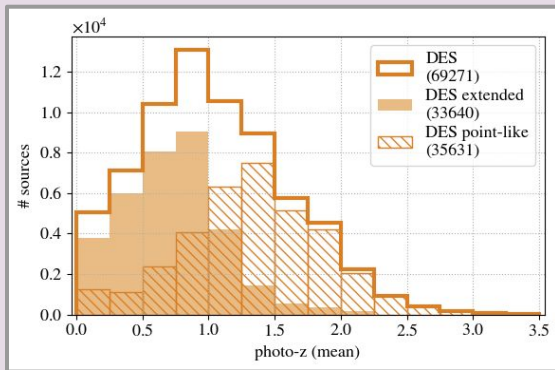
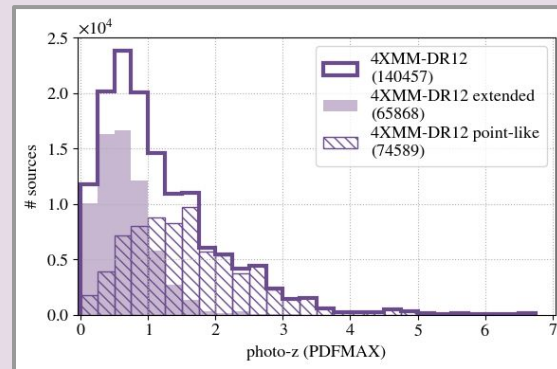
4XMM-DR12
~1000 deg²

4XMM-DR11
DES-VHS
~200 deg²

TPZ photo-z

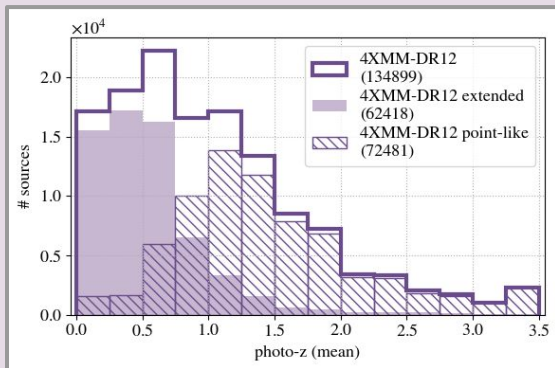


LePhare photo-z

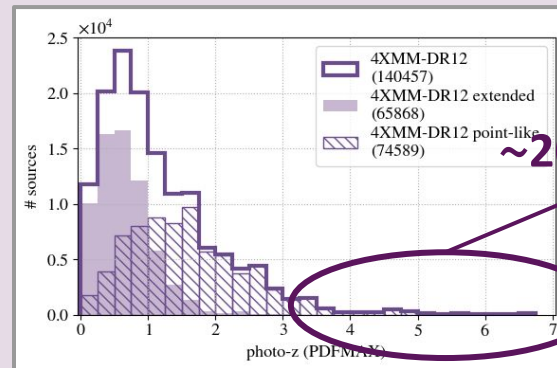


4XMM-DR12
~1000 deg²

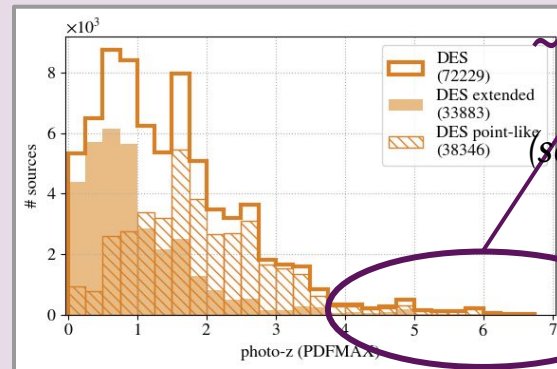
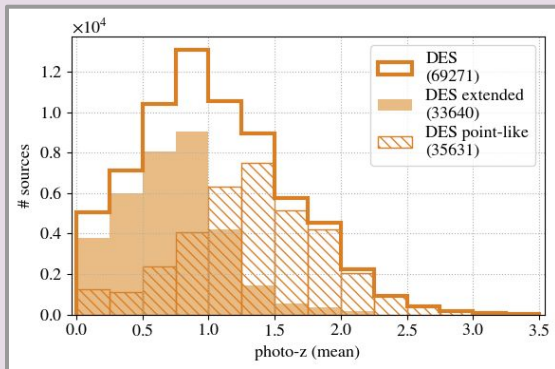
TPZ photo-z



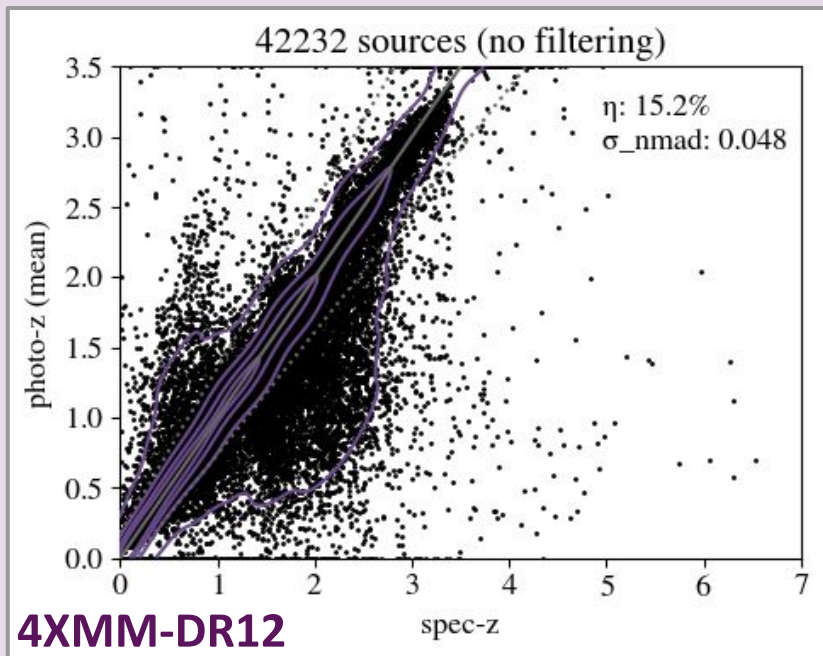
LePhare photo-z



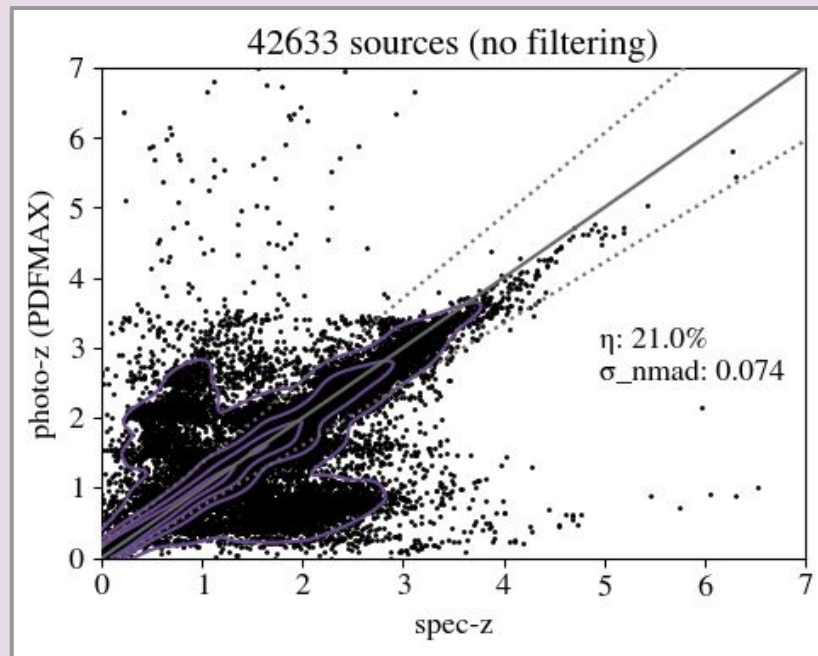
4XMM-DR11
DES-VHS
~200 deg²



TPZ photo-z

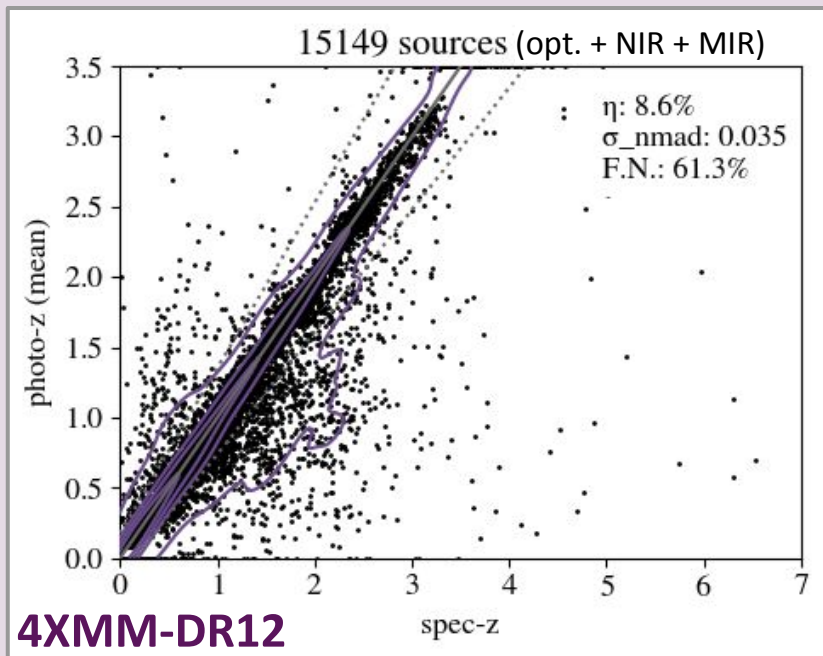


LePhare photo-z

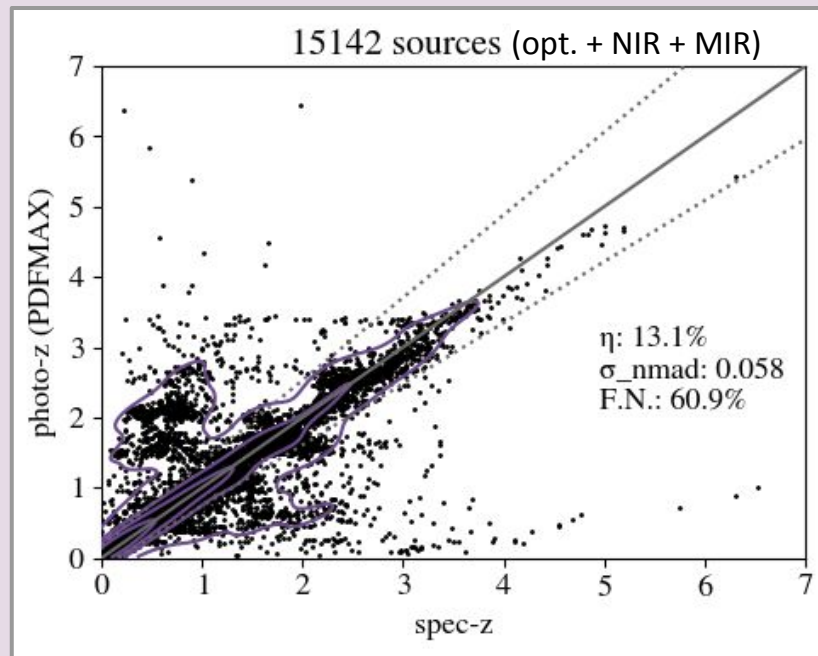


Crossmatch with with MORX spectroscopic sample

TPZ photo-z

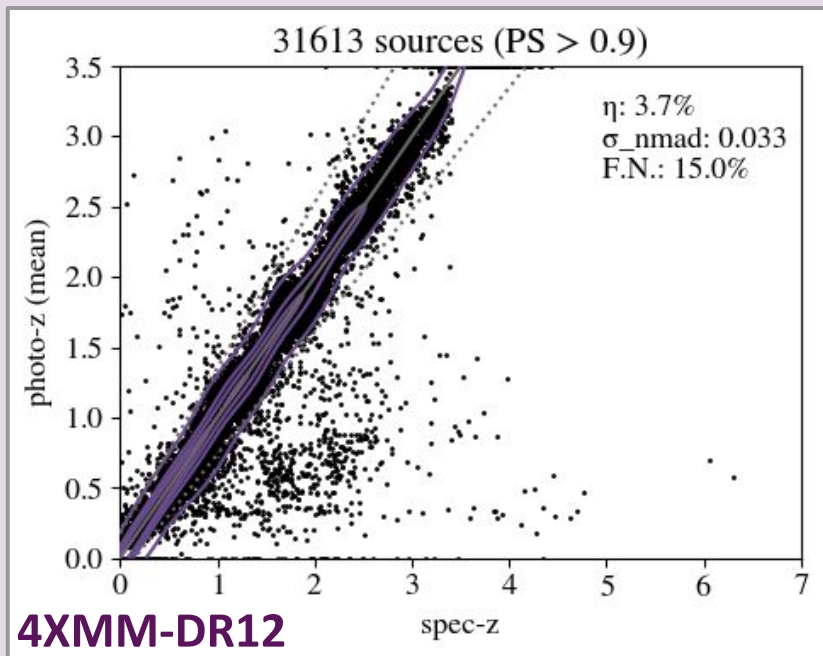


LePhare photo-z



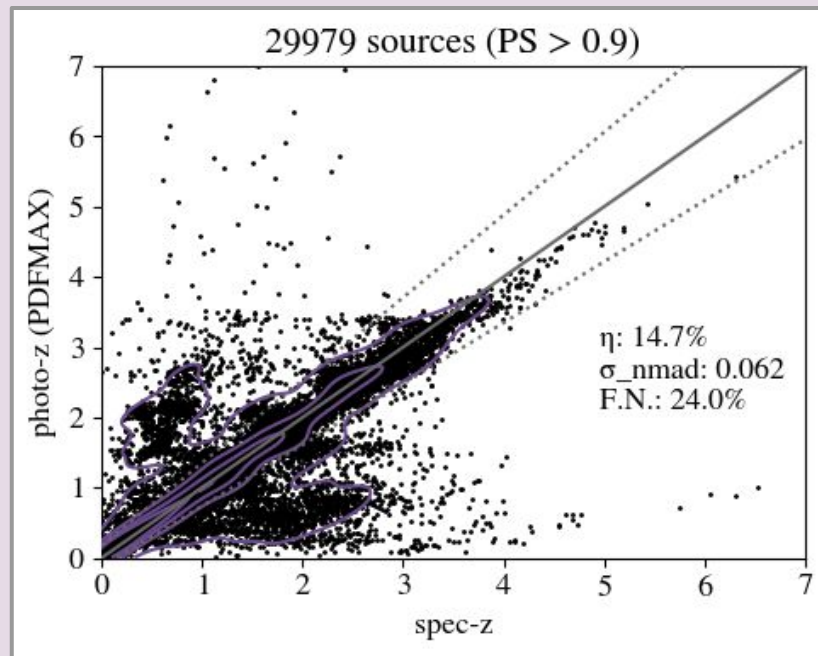
Crossmatch with MORX spectroscopic sample

TPZ photo-z



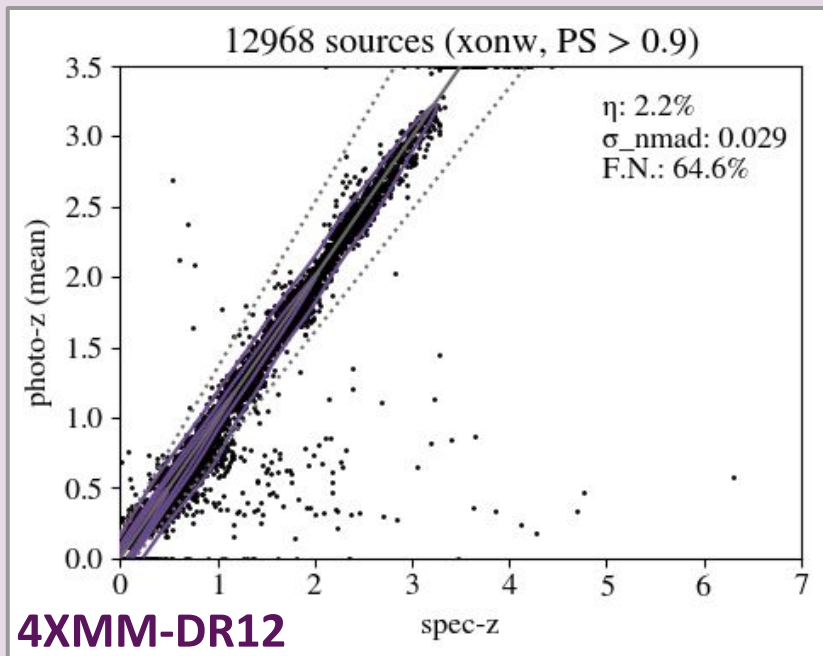
4XMM-DR12

LePhare photo-z



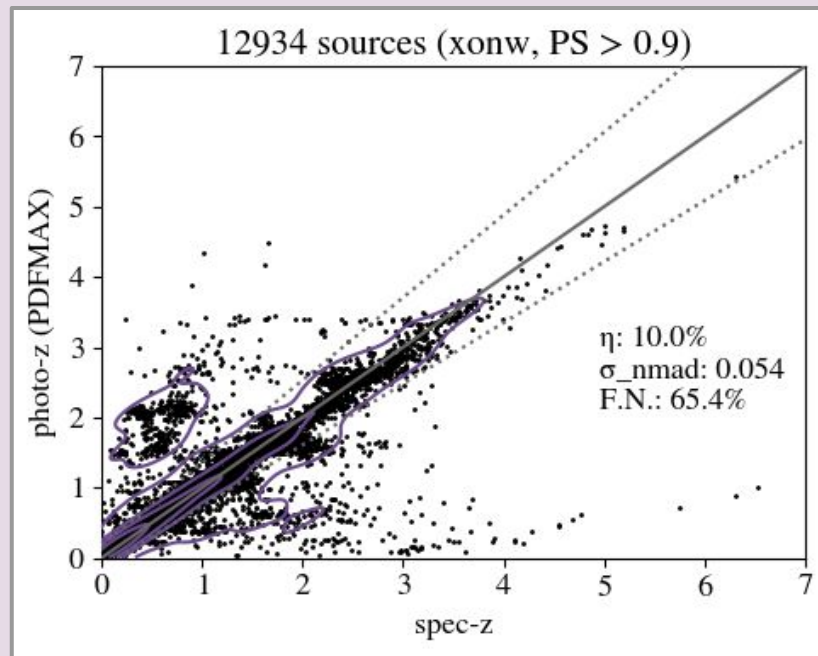
Crossmatch with MORX spectroscopic sample

TPZ photo-z



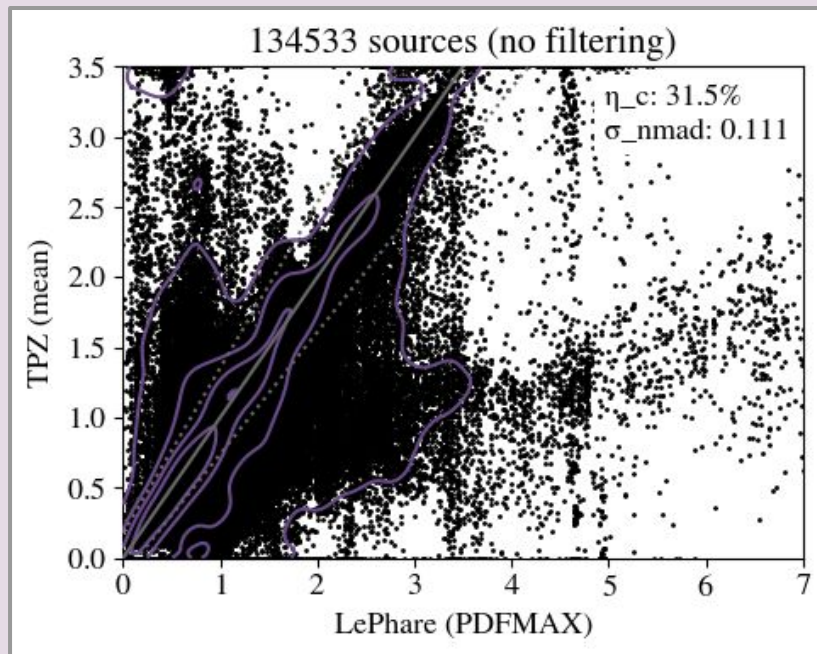
4XMM-DR12

LePhare photo-z



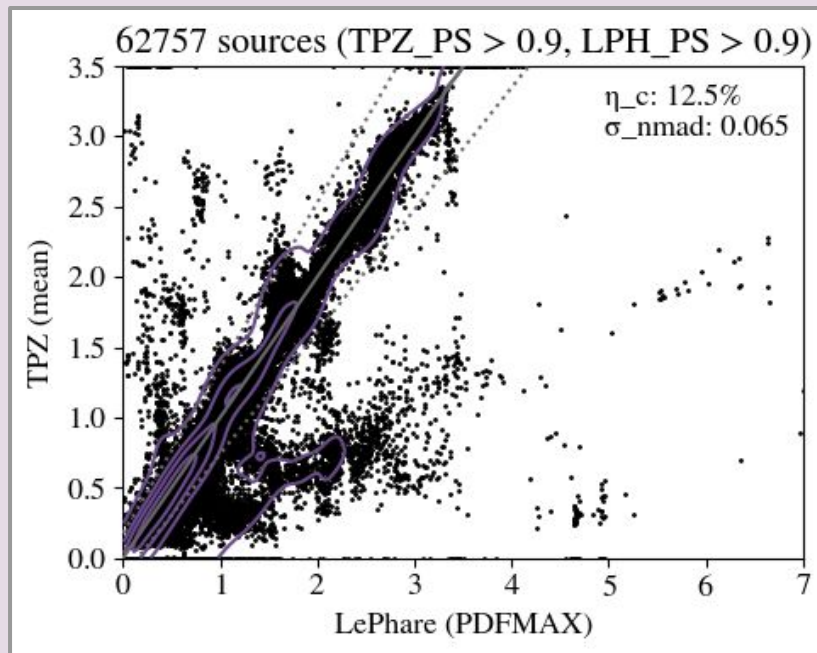
Crossmatch with MORX spectroscopic sample

4XMM-DR12



Consistent results
for ~70% of sources

4XMM-DR12



Consistent results
for ~88% of sources

SUMMARY

- **We built a catalogue of photometric redshifts for the 4XMM:**
 - 4XMM-DR12 (full sky, ~ 1000 sq.deg): $\sim 140,000$ sources
Photo-z for AGN/QSO with SDSS/PanSTARRS/SkyMapper counterparts
 - 4XMM-DR11 (DES-VHS footprint, ~ 200 sq.deg.): $\sim 72,000$ sources
Photo-z for AGN/QSO with DES counterparts
- **We provide photo-z estimations using two methods:**
 - **TPZ** (machine learning)
 - **LePhare** (template fitting)
- **The catalogue contains quality flags allowing the user to select clean samples with reliable photo-z.**
- **Improvements for the final version:**
 - Catalogue for 4XMM-DR13.
 - Include (at least partially) extragalactic sources within the galactic plane.
 - Improve LePhare photo-z for point-like (AGN dominated) sources.
 - Estimate LePhare photo-z for sources with GAIA, APASS, SUSS counterparts.
 - Better understanding of differences between TPZ and LePhare results.