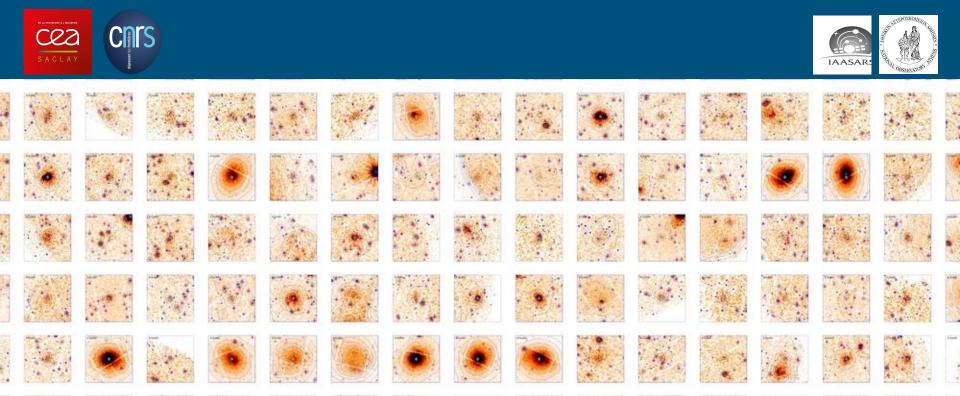
The X-CLASS survey: A catalogue of 1646 X-ray-selected galaxy clusters up to z~1.5

E. Koulouridis, N. Clerc, T. Sadibekova + the XCLASS collaboration (2021)



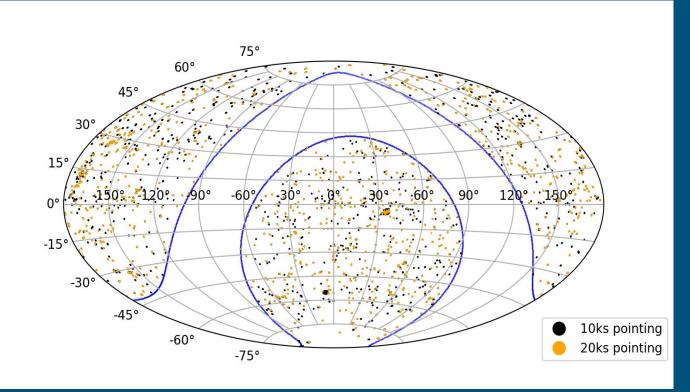
The X-CLASS collaboration

- A 20+ member international team with broad expertise.
- Europe: France, Greece, ESA scientists, Germany & UK
- International: Egypt, Uzbekistan, Japan
- Clusters, AGN multi-wavelength science + machine learning, database
- Catalogues fully public and accessible through CDS and database
- MoU regulates internal collaboration work
- Collaboration is project-driven with (currently) best-effort strategy
- Funding to be secured
- Public website in construction...

9333 XMM-Newton observations until 08/2015

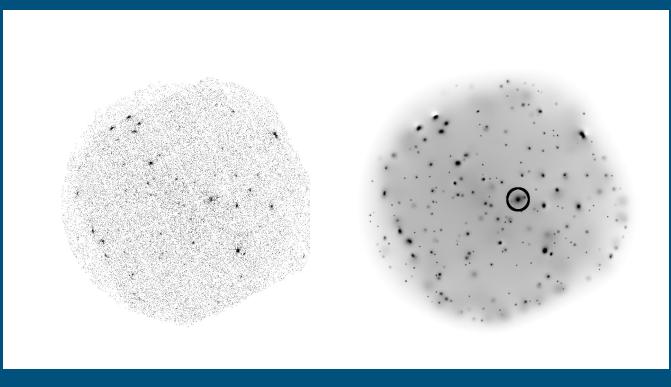
2461 observations after a cut at 10 or 20 ks in order to facilitate the computation of the selection function

total area of 269 sq. degrees

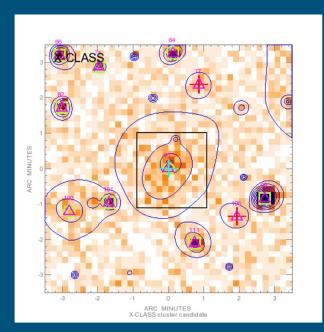


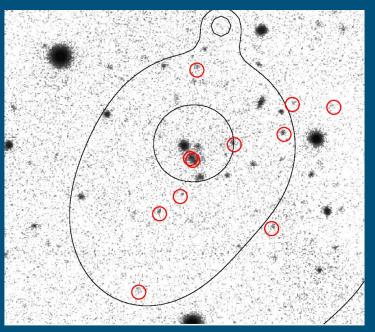
Pipeline source detection - XAmin 3.5

- 1 Creating the images in soft [0.5-2] keV band and filtering with a wavelet algorithm (Starck & Pierre 1998);
- 2 Centroid, extent and count-rate provided by SExtractor;
- 3 Maximum likelihood fitting considering PSF and beta-model (Cavaliere & Fusco-Femiano 1978). The sources then are characterized by values of their extent likelihood and extent (best-fit core radius)



Redshift validation





Cluster Xclass0219 at z=0.791 as confirmed by 11 member galaxies. Left panel: X-ray image and contours. Green circles (squares) mark detections of extended (point-like) sources as classified by the XAmin pipeline. Straight lines that cross the image are CCD gaps of the XMM-Newton detector. Right panel: i-band optical image from PanSTARRS over-plotted with X-ray contours. Red circles mark the member galaxies with available spectroscopic redshift.

Redshift validation

<u>Confirmed</u>: if three or more galaxies with concordant spectroscopic redshifts are found within the 500 kpc radius from the centre of the X-ray detection, or alternatively, if the spectrum of the brightest cluster galaxy (BCG) is available.

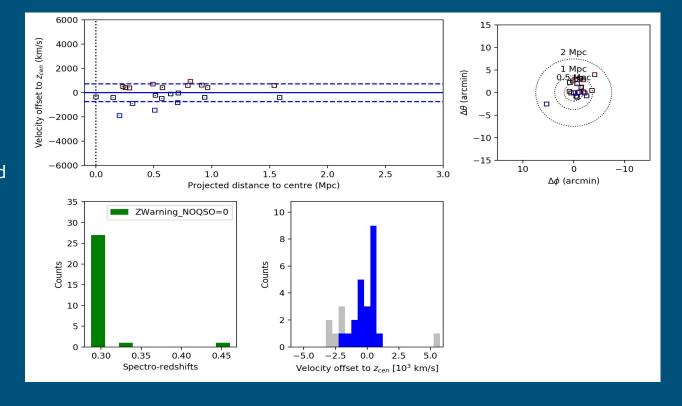
Tentative: if one or two galaxies with concordant spectroscopic redshifts are found within the 500 kpc radius.

<u>Photometric</u>: if only photometric redshift information is available in the literature or from our previous dedicated follow-up (Ridl et al. 2017).

<u>Provisional</u>: for cases where the available information does not allow us to verify the existence of a galaxy cluster in this position. Further follow-up observations are needed to safely classify these sources. Although these sources are part of the X-CLASS catalogue, they are not included in the on-line database.

SPIDERS:

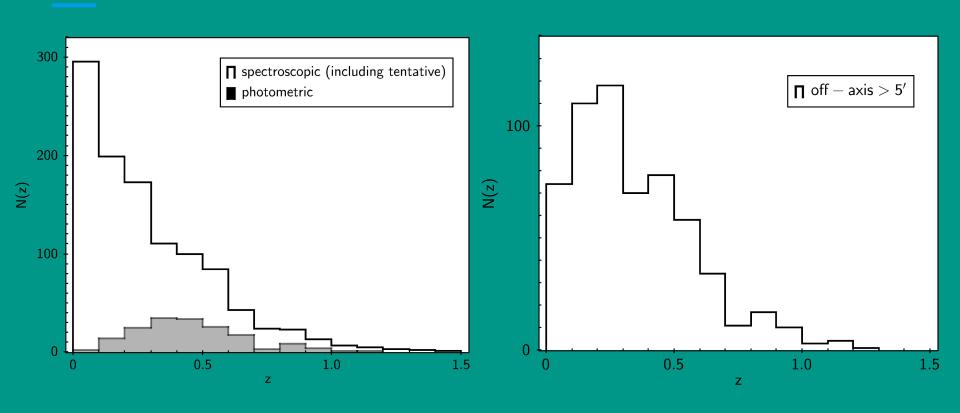
The sample was compiled based on the correlation of X-ray sources from X-CLASS with the RedMapper optical cluster catalogue, as described in (Sadibekova et al. 2014) and (Clerc et al. 2016).

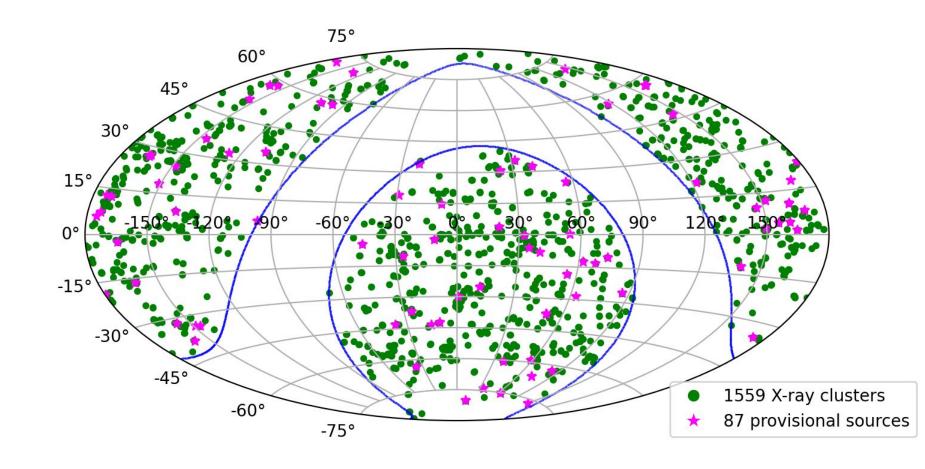


The (BOSS) spectrograph mounted on the SDSS-2.5m telescope at Apache Point Observatory (Gunn et al. 2006) was used. The catalogue contains **124 validated clusters** (out of the 142 targeted) with SPIDERS follow-up spectroscopy up to a redshift of z~0.6. The program led to the collection of **1134 spectra in X-CLASS red sequences**, with a redshift success rate approaching 99% (Clerc et al. 2020).

redshift distribution

The redshift distribution peaks at $z\sim0.1$, while if we remove the pointed observations it peaks at $z\sim0.3$.





Follow-up on-going programs

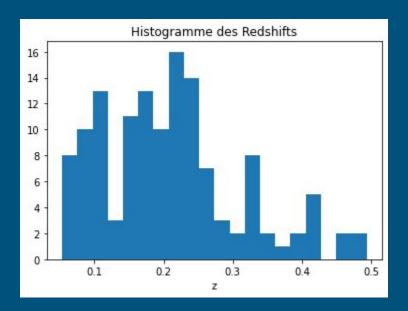
Optical identification with recent surveys (photometry, zphots, redshifts)

DESI Legacy, PanStarrs, SDSS ...

Complementary spectroscopic observations with MISTRAL@OHP

16 nights in 2021 – 2023 ; 47 new groups/clusters with redshift, mostly low-z clusters (zMED \sim 0.21)

Update the database with recent other spectroscopic measurements



Follow-up on-going programs

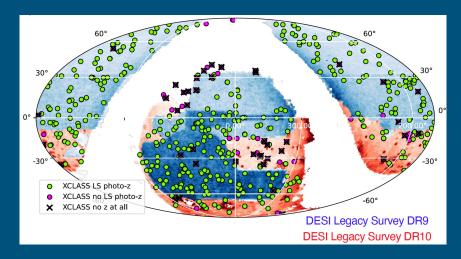
Optical identification with recent surveys (photometry, zphots, redshifts): DESI Legacy, PanStarrs, SDSS ...

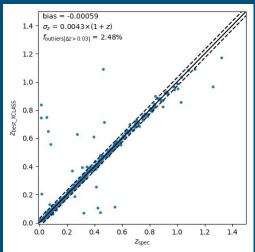
Previous multi-λ photometric follow-up with GROND@2.2m – ESO (Ridl et al. 2017). grizJHK photometry, 400+ targets.

~230 clusters already published with zMED = 0.39

Still 100+ (distant) clusters remain to analyze

For Southern targets available with DESI, definition of the most probable redshift from the photo-z identifications





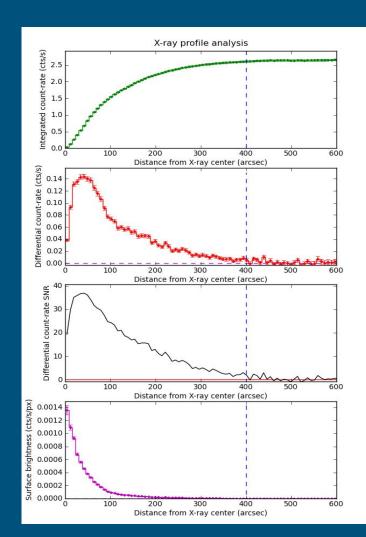
XCLASS cluster photo-z vs. Spectro-z validation (Sarron et al.)

Interactive flux measurements

Interactive (manual) mode enabling the user to:

- refine the X-ray cluster centre,
- remove or correct areas incorrectly masked by XAmin (CCD gaps, unresolved blended sources, FOV edge cases),
- re-estimate the background level according to the cluster brightness and extension to get a more precise count-rate measurement
- optimise the measurements in cases where the source is detected on the missing part of MOS1, and
- set a more accurate and reliable value for the source radii when the growth curve algorithm has failed because of background overestimation (field source contamination, missing part of MOS1, edge effects).

the count rates are computed in six different bands, namely [0.5-2], [2-0], [0.5-0.9], [1.3-2], [2-5] and [5-7] keV, using a full exposure to obtain the highest signal-to-noise ratio.



Cluster selection criteria

	Ar	rea	
r.a. min dec min		r.a. max dec max	help
	clear	get	load save

xclass min redshift min status min		~	xclass max redshift max status max	· ·
	clear		get	load save

off radius max detection ML max extent ml max	
detection ML max	
extent ml max	
extent max	
counts max	
rate max	
flux max	
	rate max

Selection					
clear get load save all criteria. <u>Refresh</u> page. select 1559 clusters selected <u>Display</u> the selected subset of clusters.					

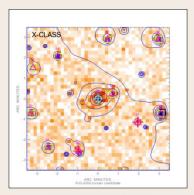
More information is available in the user's guide



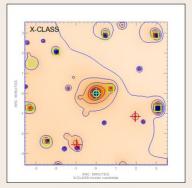
The new X-CLASS cluster catalogue

Subset of clusters Order by xclass then by display obs redshift total rate xclass R.A. pipeline Dec pipeline R.A. measured Dec measured NED status profile 193.4380 10.1954 193.4380 10.1951 0020 0001930301 10ks 0.654 confirmed 0.049 data 0023 194.2860 -17.4119194 2920 -17.40640010420201 10ks 0.047 confirmed 3 738 data 0033 193,6790 -29.2227 193,6740 -29.2230 0030140101 10ks 0.056 confirmed 5.882 data 0034 193,5950 -29.0162 193,5930 0030140101 10ks 0.053 confirmed 4.362 -29.0131 data 0035 196.2740 -10.2802196.2740 -10.27870032141201 10ks 0.34 photometric 0.047 data 36.5674 -2.6651 36.5677 0038 -2.66630037981801 10ks 0.056 confirmed 0.167 data 0039 36.4987 -2.827236,4990 -2.82750037981801 10ks 0.281 confirmed 0.033 data 0040 35.1871 -3.4339 35.1886 -3.4339 0037982601 10ks 0.327 confirmed 0.050 data 0042 150.1230 -19.6282 150.1220 -19.6292 0041180301 10ks no redshift 0.057 data 0044 202.4460 11 6835 202 4490 11.6848 0041180801 10ks 0.204 confirmed 0.087 data 0047 172.9830 -19.9229 172.9800 -19.9271 0042341001 10ks 0.307 confirmed 3.254 data 0048 173.0280 -19.8611 173.0280 -19.8614 0042341001 10ks 0.307 confirmed 0.154 data 0050 172.8110 -19.9326172.8130 -19.93430042341001 10ks 0.46 photometric 0.025 data 0051 177.6130 1.7580 177,6160 1.7580 0044740201 10ks no redshift. qo 0.036 data 0054 145.9370 16.7402 145.9380 16.7381 0046940401 10ks 0.180 confirmed 0.136 data 0056 145.8820 16.6656 145.8860 16.6671 0046940401 10ks 0.255 confirmed 0.202 data 0057 145,9920 16.6871 145.9950 16.6875 0046940401 10ks 0.253 confirmed 0.057 data 31.9565 2.1553 31.9576 0059 2.1567 0052140301 20ks 99.90 no redshift 0.042 data 0062 44.1414 0.1037 44.1417 0.1033 0056020301 10ks 0.362 confirmed 0.778 data 0065 339.2510 -15.2730339.2520 -15.27310056021601 10ks 0.31 photometric 0.316 data 0075 10.4501 -9.4575 10.4507 -9.4569 0723802201 20ks 0.056 confirmed 15.529 data 0078 10.7223 -9.5697 10.7225 -9.5701 0065140201 10ks 0.41 photometric 0.101 data 10.5231 0079 10.5228 -9.6026 -9.6029 0065140201 10ks 0.055 tentative 0.033 data go 39,4926 -52.3934 39,4929 0082 -52.3937 0067190101 10ks 0.136 confirmed 0.259 data 0083 148,4240 1.6999 148.4240 1.6995 0070940401 10ks 0.097 confirmed 0.796 data go 348.7650 0086 -58.9351 348.7660 -58.9354 0081340301 10ks photometric 0.034 data

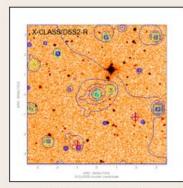
pipeline measurements and products



Raw X-ray image + wavelet contours



Wavelet image + wavelet contours



DSS2 band image + wavelet contours

Green circle : Xamin extended source (C1 or C2)
Magenta

triangle : SExtractor detection

Red cross : Point-like source having a detection likelihood < 15

Cyan cross: position of the cluster candidate center Green square: Xamin point source

Zoom optical image (4'x4')

Total (PN + MOS1 + MOS2) Xray Properties -

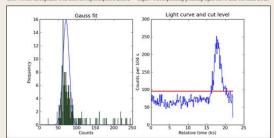
Off-axis radius:	4.4	arcmin
ML of detection:	360.664	
ML of extension:	99.1882	
Xamin extension:	13.6669	arcsec
Total counts:	298.931	total (PN + M1 + M2) observed in band B2 (0.5-2 keV) counts
Total rate:	0.049222	total (PN + M1 + M2) observed in band B2 (0.5-2 keV) count rate

PN

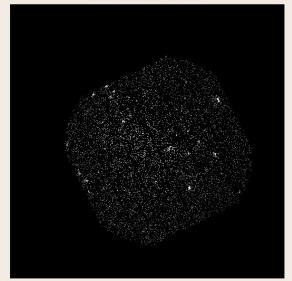
Cut: 71.00 Effective time: 16185. (s)

Pilter : Thin1 Mode : PrimePullWindow

Left : event histograms with best fit superimposed (blue). Right : corresponding pointing light curve and cut limit (red).



Raw image of events in the [0.5-2] keV energy range.



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Link to NED and search within 3' radius

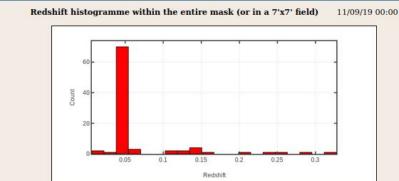


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Redshift validation

Redshift value :	0.047	error +	-
Velocity dispersion :		km/s error +	-
Number of galaxies :	63		
Final status :	confirmed	_	
Last update :	27/11/20 00:00	Set by : sadibekova	1
194	axies in the cluster .27892 -17.41672 .37625 -17.45736 .34717 -17.27622	9.041649 8.042489 9.0422876	



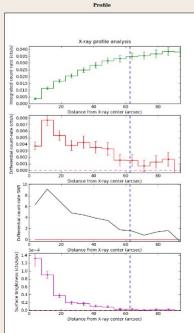
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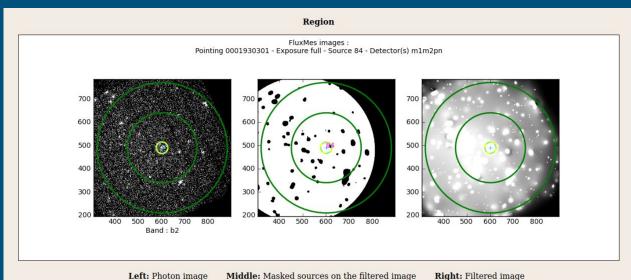
xclass 0020

CR1	b2 [0.5-2] keV	0.0344	error ±	0.0025
CR2	b3 [2-10] keV	-0.0006	error ±	0.0022
CR3	b21 [0.5-0.9] keV	0.0160	error ±	0.0018
CR4	b24 [1.3-2] keV	0.0099	error ±	0.0013
CR5	b31 [2-5] keV	0.0028	error ±	0.0014
CR6	b33 [5-7] keV	0.0002	error ±	0.0010

Rfit	62.5	arcsec	
N _H	1.4	* 10 ²⁰ cm ⁻²	



Interactive flux measurements



- · Dash cyan circle: limit of the automatic integration (not
- Yellow circle: limit of the interactive integration
 Green circles: background region

The X-CLASS Cluster Survey

- The X-CLASS survey: A catalogue of 1646 X-ray-selected galaxy clusters up to z~1.5 Koulouridis et al., 2021, A&A, 652A, 12K
- Multiwavelength classification of X-ray selected galaxy cluster candidates using convolutional neural networks Kosiba et al. 2020, MNRAS. 496. 4141K
- Cosmology with XMM galaxy clusters: the X-CLASS/GROND catalogue and photometric redshifts Ridl et al., 2017, MNRAS, 468, 662
- SPIDERS: the spectroscopic follow-up of X-ray-selected clusters of galaxies in SDSS-IV Clerc et al., 2016, MNRAS, 463, 4490
- The X-CLASS-redMaPPer galaxy cluster comparison Sadibekova et al., 2014, A&A, 571, 87
- The cosmological analysis of X-ray cluster surveys II. Application of the CR–HR method to the XMM archive Clerc et al., 2012, MNRAS, 423, 3561

DATABASE: https://xmm-xclass.in2p3.fr contact: ekoulouridis@noa.gr

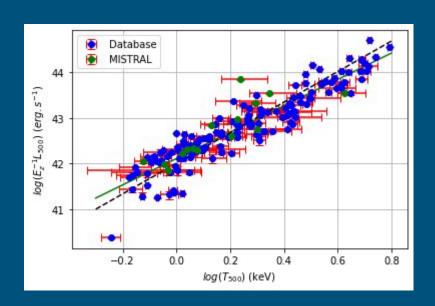
What next?

Importance of the redshift measurements to study the physical properties of clusters: X-ray profile and luminosity, temperature, R500, optical properties of groups and clusters

Specificity of the XCLASS sample : dominated by faint X-ray sources, mostly low redshift groups and poor clusters + high-redshift clusters

Constraints of the LX – T – M relation, then the mass function at $M < 5 \times 1013 \text{ Msol}$

A cosmological analysis is ongoing, specific to « small clusters »!



The X-CLASS z<0.2 luminosity-temperature relation (Moysan et al. in prep.)

Thank you!